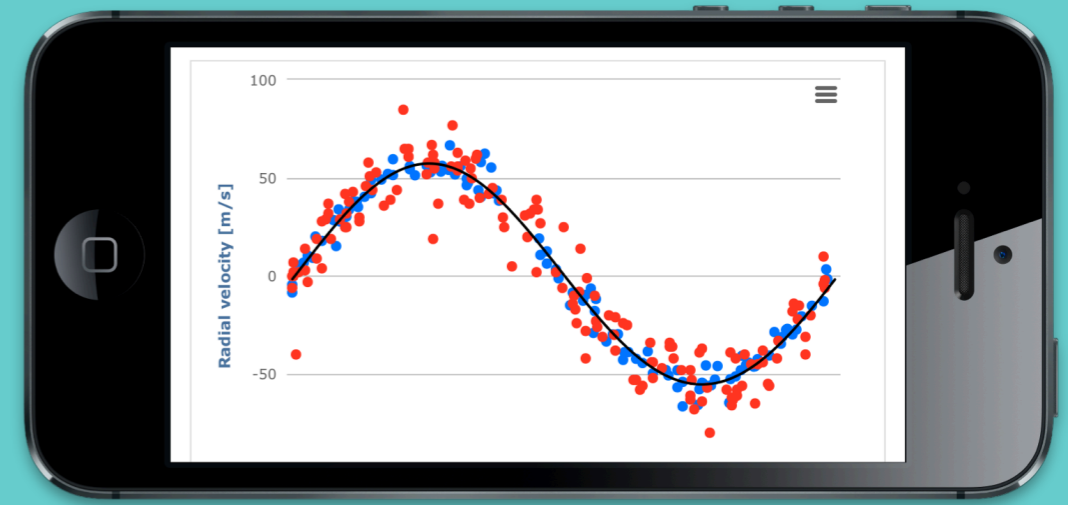
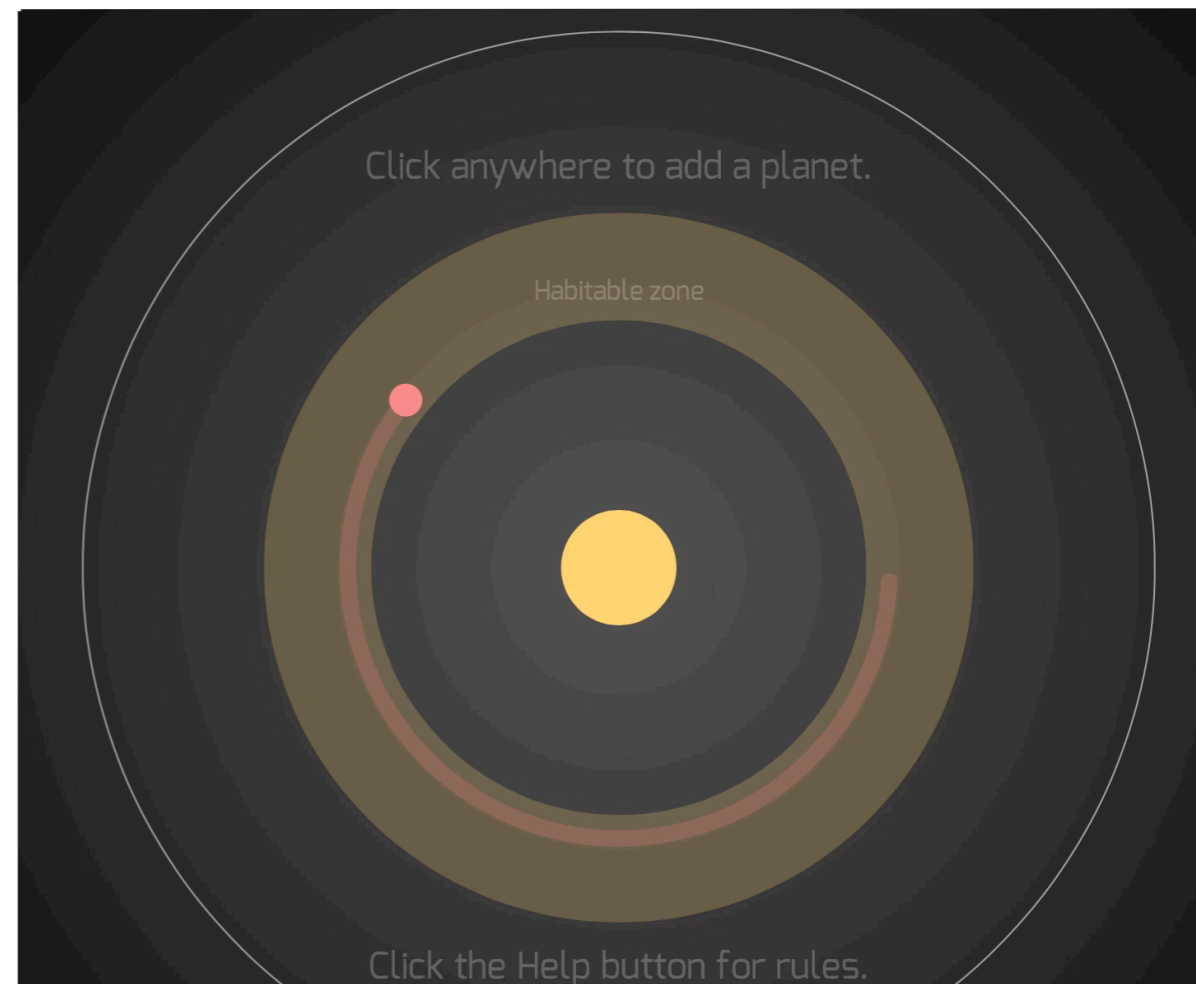
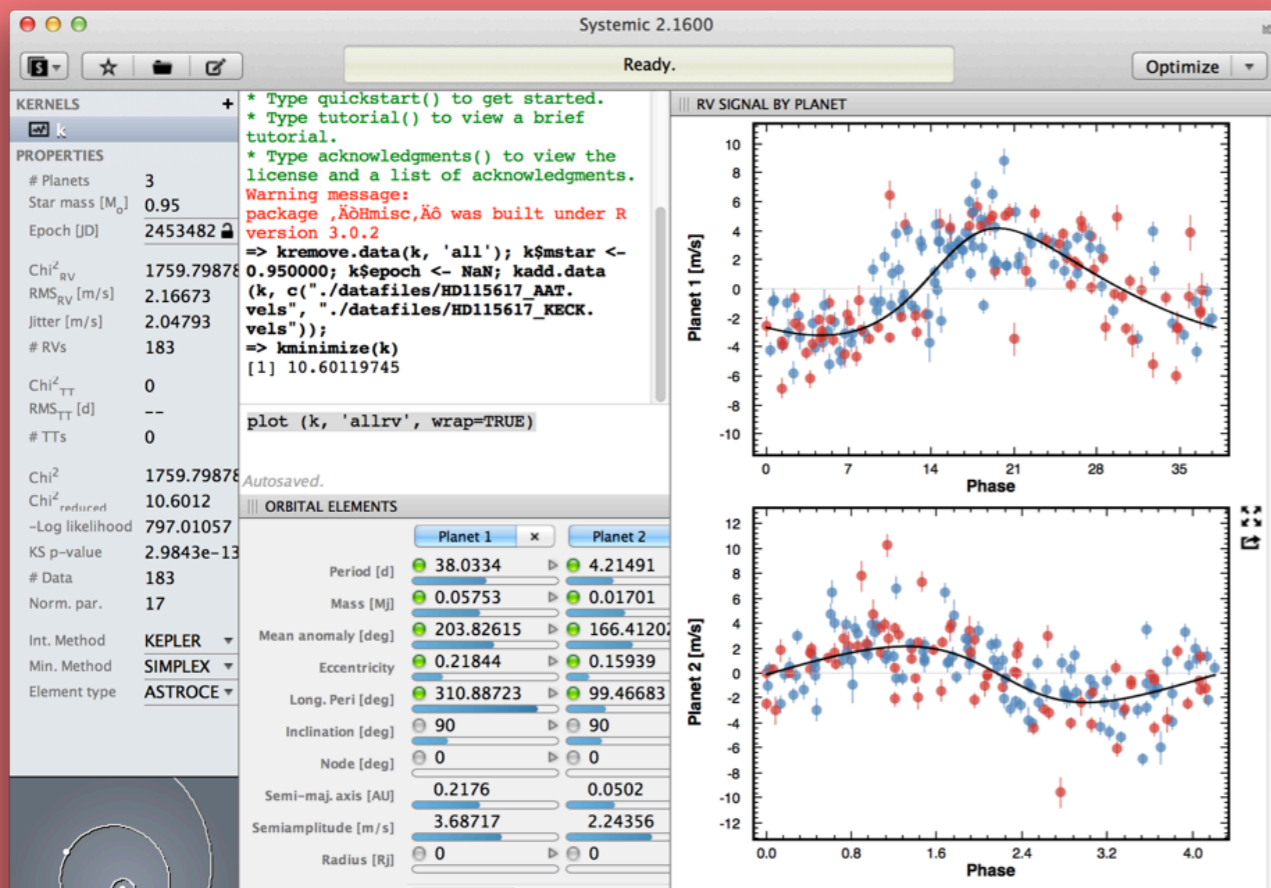


You, Too, Can Make Spiffy Online Web Apps for Outreach and \$\$\$

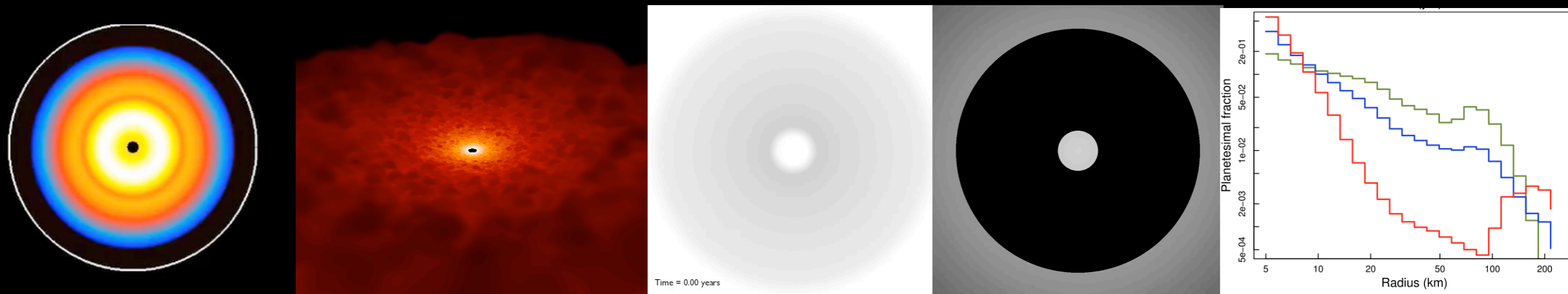


Stefano Meschiari
UT Austin, W.J. McDonald Fellow

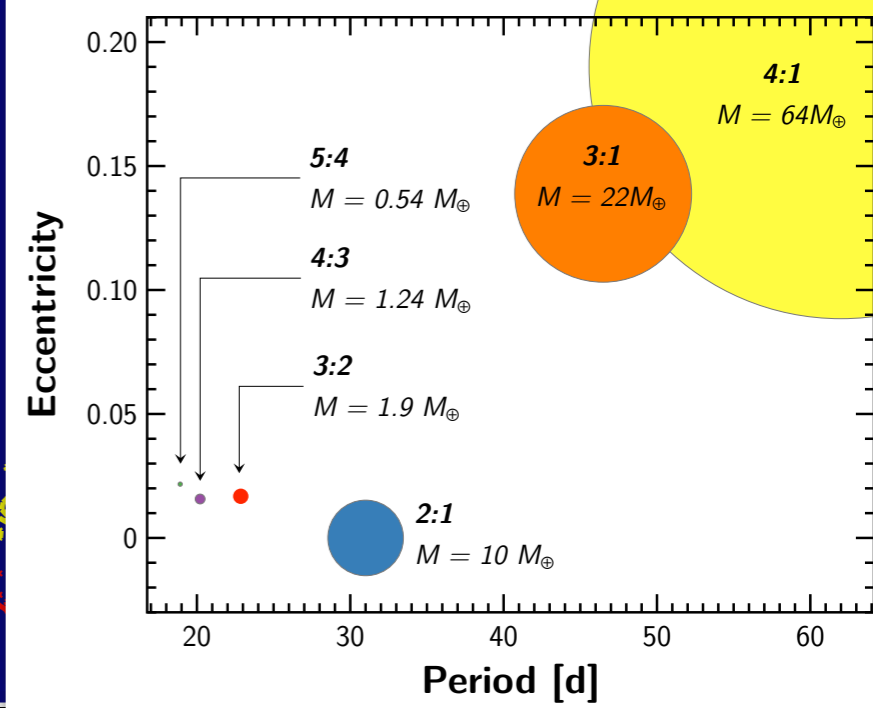
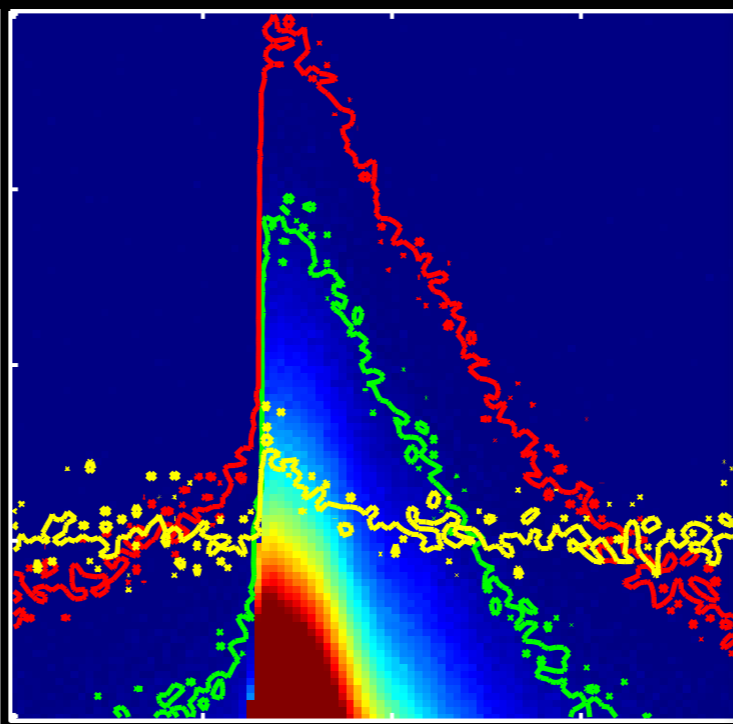
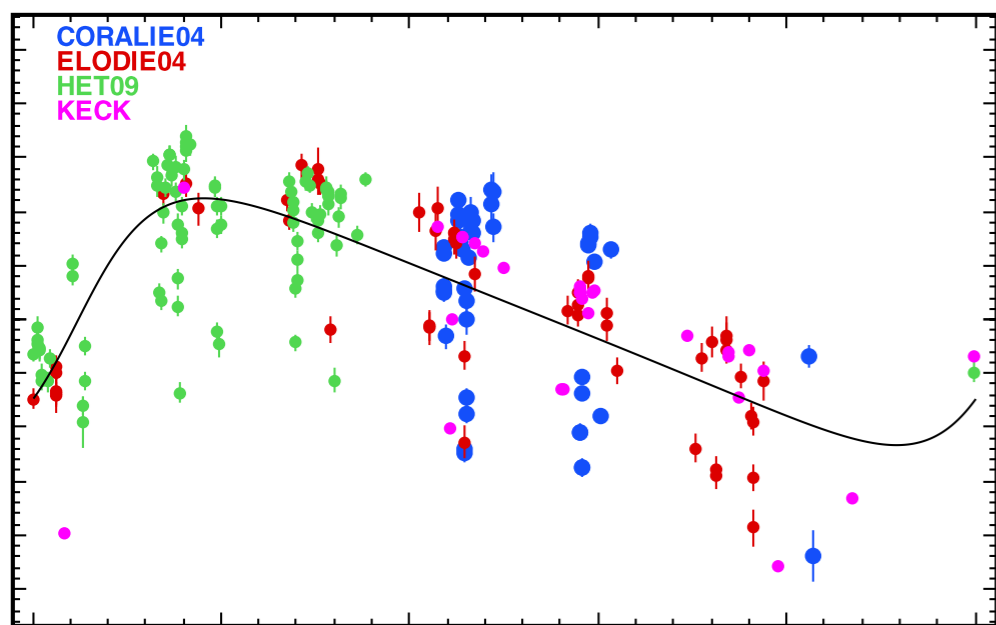
UT Austin GSPS
March 21, 2014



If you don't know who I am:



My science shtick is planet formation & exoplanet detection.



I'm also very interested in outreach, especially when it doesn't involve me physically standing in front of an audience.

Shy panda



Systemic:

One Software Package to Rule Them All

What is Systemic?

Systemic is an open-source software package for analyzing and modelling exoplanetary time series (primarily Radial Velocities and transit timing)

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What is Systemic?

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What is “All”?

Systemic:

One Software Package to Rule Them All

What is Systemic?

Systemic is an open-source software package for analyzing and modelling exoplanetary time series (primarily Radial Velocities and transit timing)

What is “All”?

● **Science**

● **Teaching & outreach**

● **A fun treat**

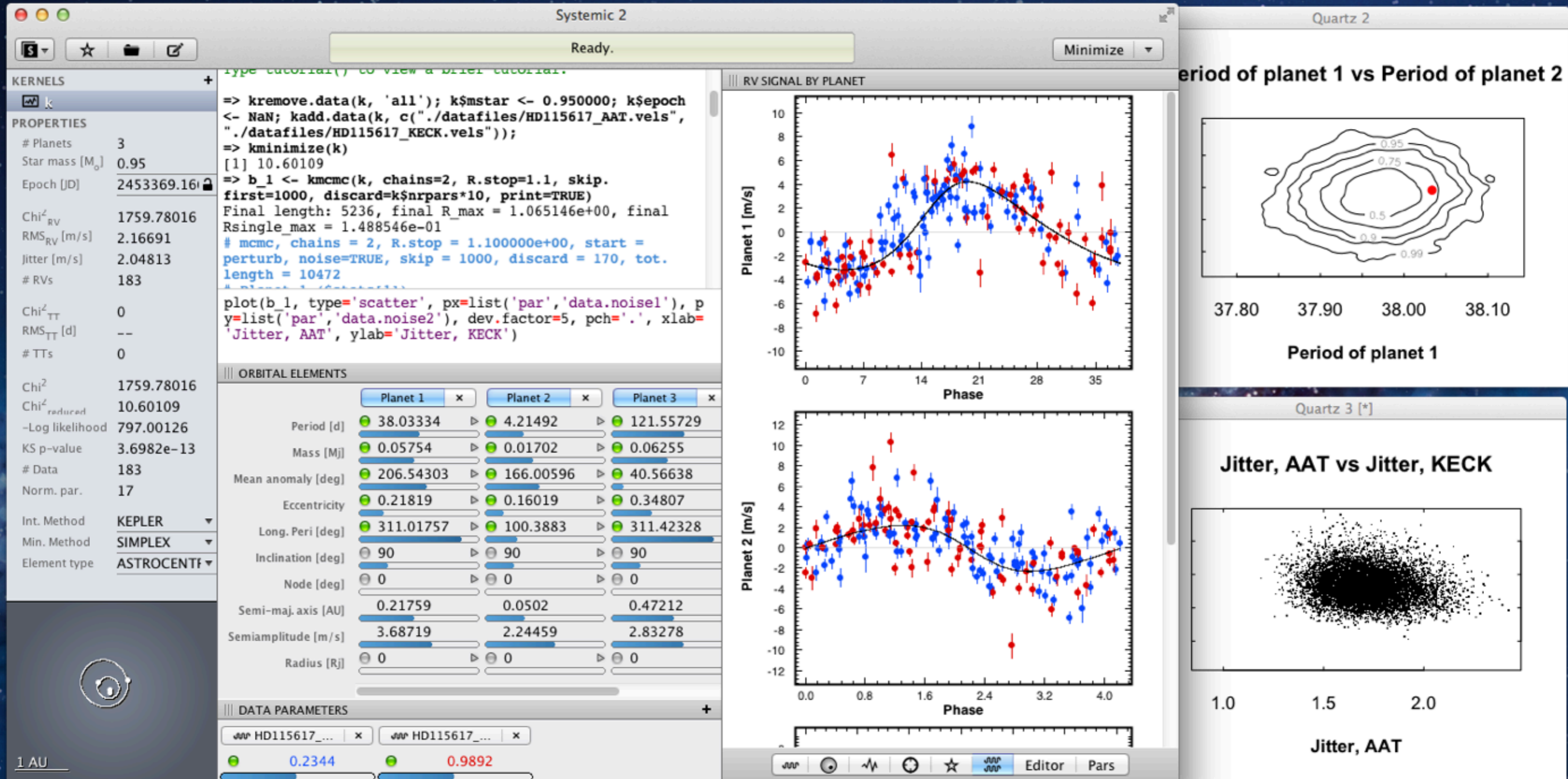
Science

Collaborators:

Greg Laughlin, Russell Hanson,
Jenn Burt, Steve Vogt (UCSC),
Paul Butler (Carnegie), Joel Green (UT)

Systemic 2

<http://www.stefanom.org/systemic>



Meschiari+ '09, '10, '11, Meschiari+ '14 (in prep.)

Systemic 2

Ready.

KERNELS

PROPERTIES

- # Planets: 3
- Star mass [M_{\odot}]: 0.95
- Epoch [JD]: 2453369.16
- Chi²_{RV}: 1759.78016
- RMS_{RV} [m/s]: 2.16691
- Jitter [m/s]: 2.04813
- # RVs: 183
- Chi²_{TT}: 0
- RMS_{TT} [d]: --
- # TTs: 0
- Chi²: 1759.78016
- Chi²_{reduced}: 10.60109
- Log likelihood: 797.00126
- KS p-value: 3.6982e-13
- # Data: 183
- Norm. par.: 17
- Int. Method: KEPLER
- Min. Method: SIMPLEX
- Element type: ASTROCENTR

```

=> kremove.data(k, 'all'); k$star <- 0.950000; k$epoch
<- NaN; kadd.data(k, c("./datafiles/HD115617_AAT.vels",
"./datafiles/HD115617_KECK.vels"));
=> kminimize(k)
[1] 10.60109
=> b_1 <- kmcmc(k, chains=2, R.stop=1.1, skip.
first=1000, discard=k$nrpars*10, print=TRUE)
Final length: 5236, final R_max = 1.065146e+00, final
Rsingle_max = 1.488546e-01
# mcmc, chains = 2, R.stop = 1.100000e+00, start =
perturb, noise=TRUE, skip = 1000, discard = 170, tot.
length = 10472

plot(b_1, type='scatter', px=list('par','data.noise1'), p
y=list('par','data.noise2'), dev.factor=5, pch='.', xlab=
'Jitter, AAT', ylab='Jitter, KECK')
  
```

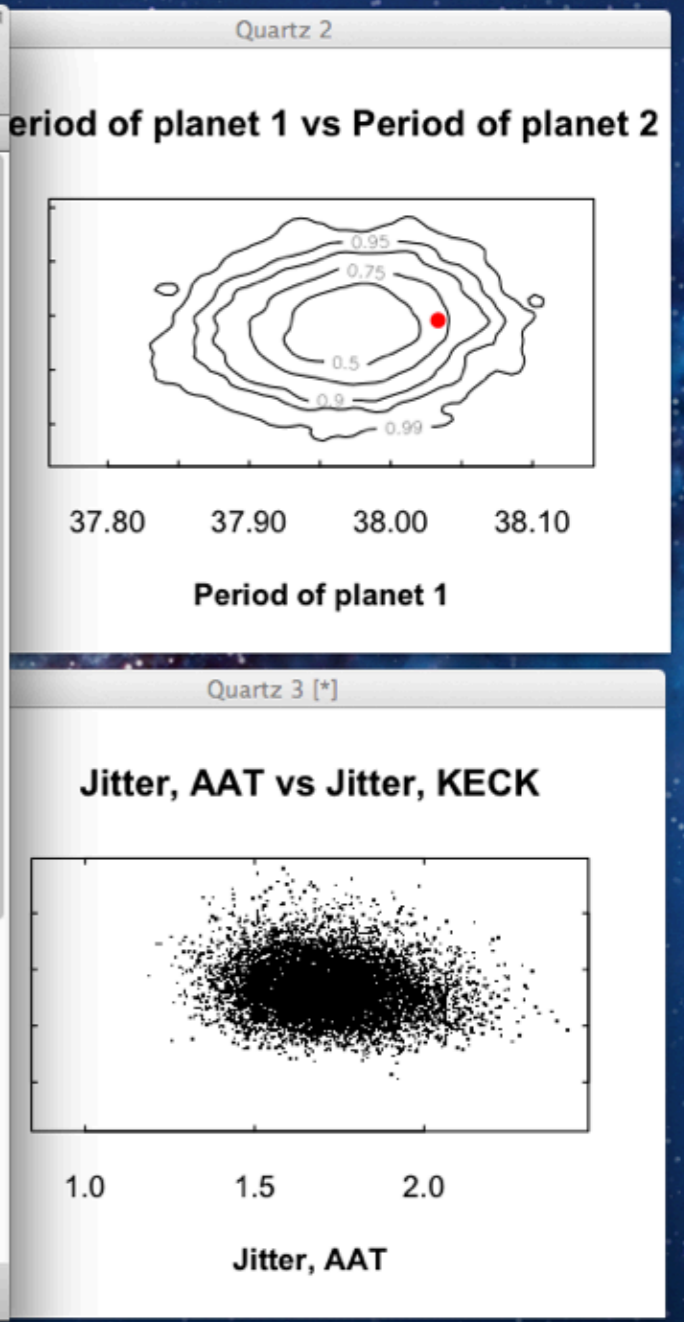
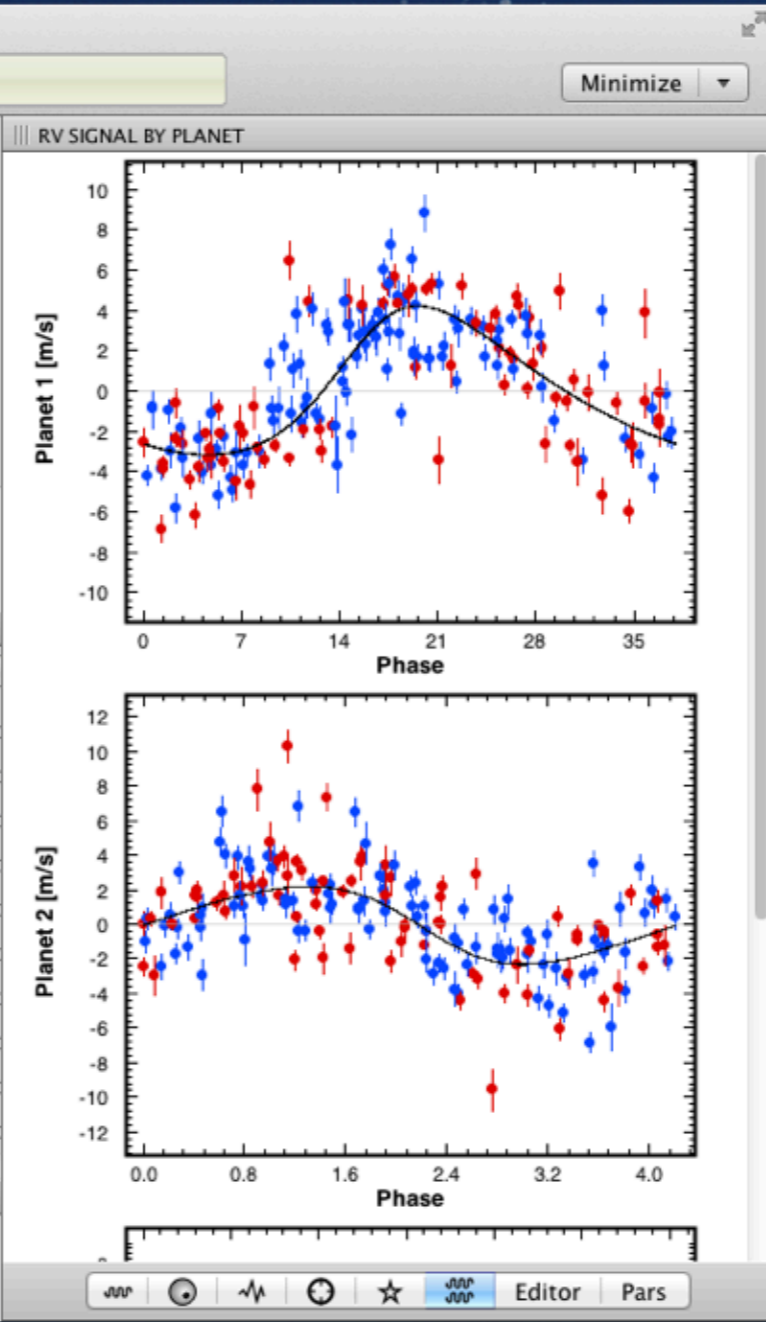
ORBITAL ELEMENTS

	Planet 1	Planet 2	Planet 3
Period [d]	38.03334	4.21492	121.55729
Mass [Mj]	0.05754	0.01702	0.06255
Mean anomaly [deg]	206.54303	166.00596	40.56638
Eccentricity	0.21819	0.16019	0.34807
Long. Peri [deg]	311.01757	100.3883	311.42328
Inclination [deg]	90	90	90
Node [deg]	0	0	0
Semi-maj. axis [AU]	0.21759	0.0502	0.47212
Semi-amplitude [m/s]	3.68719	2.24459	2.83278
Radius [Rj]	0	0	0

DATA PARAMETERS

HD115617_... 0.2344 HD115617_... 0.9892

1 AU



Model statistics (Chi², log likelihood, etc.)

Systemic 2

Ready.

KERNELS

PROPERTIES

- # Planets: 3
- Star mass [M_⊙]: 0.95
- Epoch [JD]: 2453369.16
- Chi²_{RV}: 1759.78016
- RMS_{RV} [m/s]: 2.16691
- Jitter [m/s]: 2.04813
- # RVs: 183
- Chi²_{TT}: 0
- RMS_{TT} [d]: --
- # TTs: 0
- Chi²: 1759.78016
- Chi²_{reduced}: 10.60109
- Log likelihood: 797.00126
- KS p-value: 3.6982e-13
- # Data: 183
- Norm. par.: 17
- Int. Method: KEPLER
- Min. Method: SIMPLEX
- Element type: ASTROCENTR

```

=> kremove.data(k, 'all'); k$star <- 0.950000; k$epoch
<- NaN; kadd.data(k, c("./datafiles/HD115617_AAT.vels",
"./datafiles/HD115617_KECK.vels"));
=> kminimize(k)
[1] 10.60109
=> b_1 <- kmcmc(k, chains=2, R.stop=1.1, skip.
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Final length: 5236, final R_max = 1.065146e+00, final
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# mcmc, chains = 2, R.stop = 1.100000e+00, start =
perturb, noise=TRUE, skip = 1000, discard = 170, tot.
length = 10472
plot(b_1, type='scatter', px=list('par','data.noise1'), p
y=list('par','data.noise2'), dev.factor=5, pch='.', xlab=
'Jitter, AAT', ylab='Jitter, KECK')
    
```

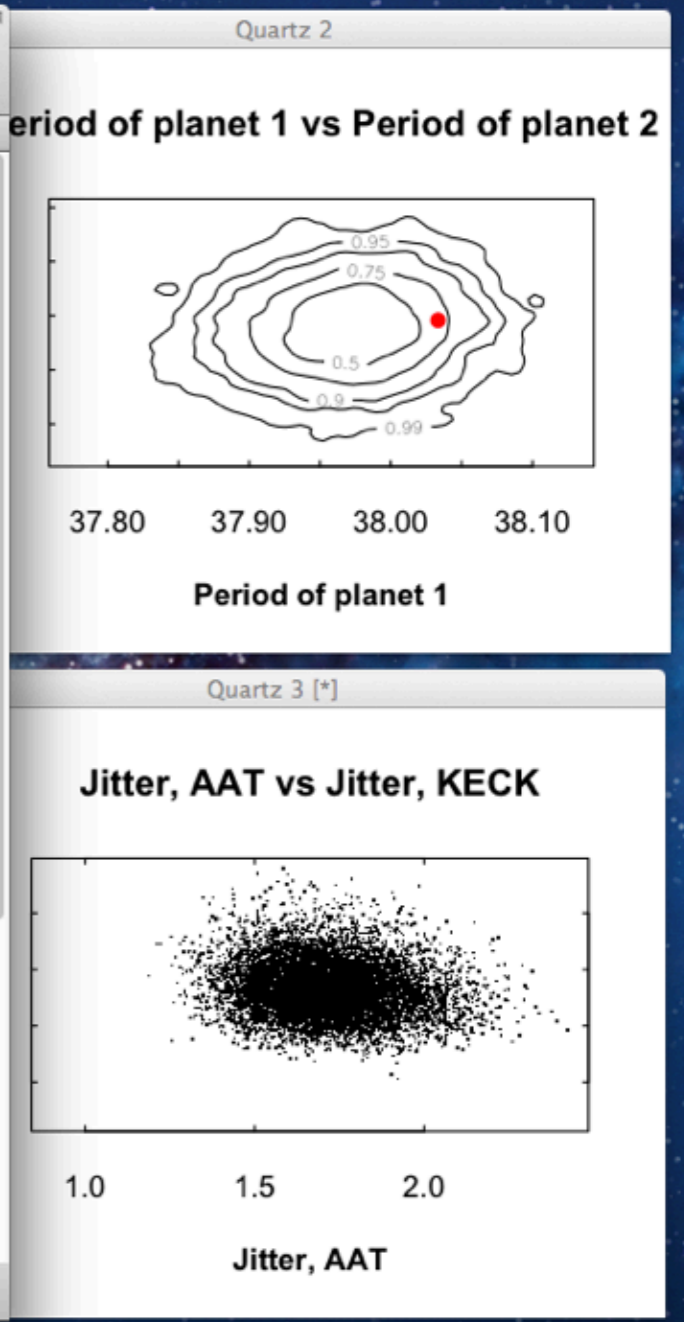
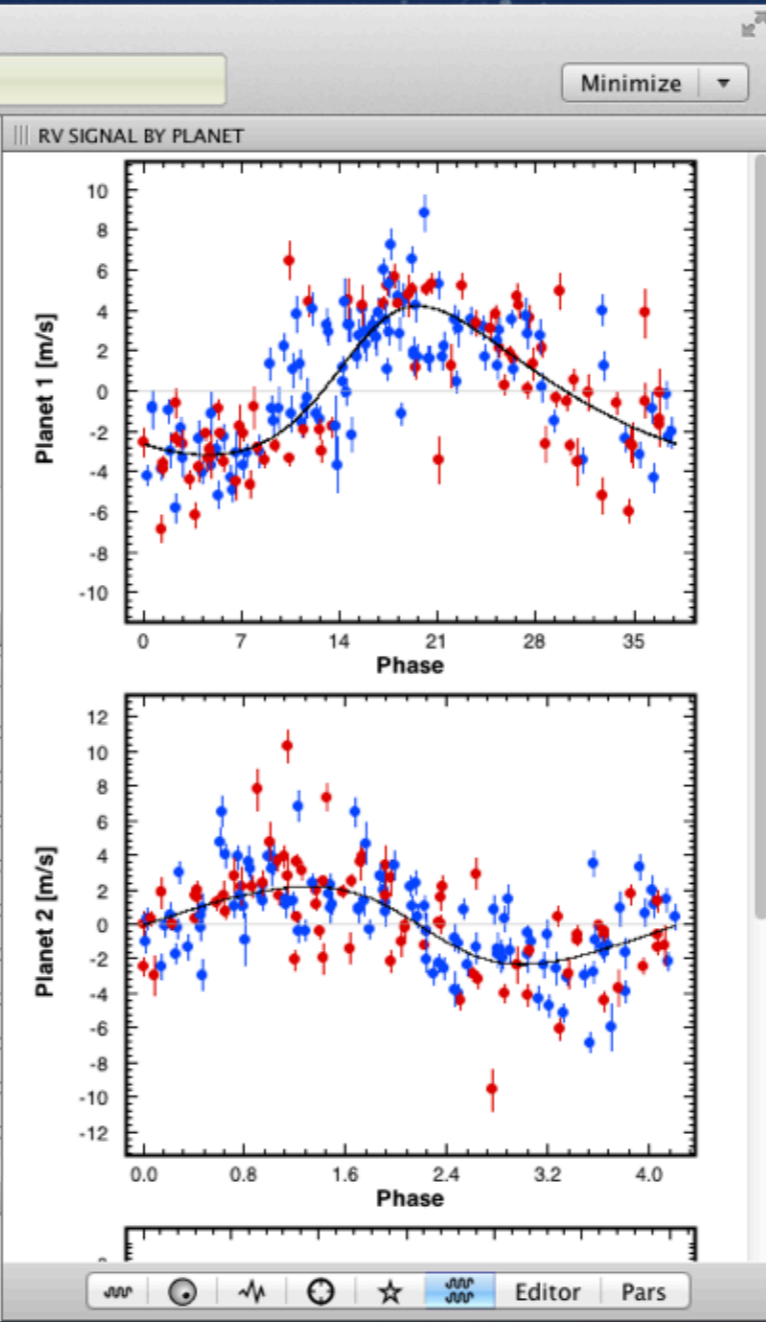
ORBITAL ELEMENTS

	Planet 1	Planet 2	Planet 3
Period [d]	38.03334	4.21492	121.55729
Mass [M _J]	0.05754	0.01702	0.06255
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Eccentricity	0.21819	0.16019	0.34807
Long. Peri [deg]	311.01757	100.3883	311.42328
Inclination [deg]	90	90	90
Node [deg]	0	0	0
Semi-maj. axis [AU]	0.21759	0.0502	0.47212
Semi-amplitude [m/s]	3.68719	2.24459	2.83278
Radius [R _J]	0	0	0

DATA PARAMETERS

HD115617_... 0.2344

HD115617_... 0.9892



Model statistics Command line

(Chi², log likelihood, etc.)

Systemic
Ready.
Minimize

KERNELS

PROPERTIES

- # Planets 3
- Star mass [M_⊙] 0.95
- Epoch [JD] 2453369.16
- Chi²_{RV} 1759.78016
- RMS_{RV} [m/s] 2.16691
- Jitter [m/s] 2.04813
- # RVs 183
- Chi²_{TT} 0
- RMS_{TT} [d] --
- # TTs 0
- Chi² 1759.78016
- Chi²_{reduced} 10.60109
- Log likelihood 797.00126
- KS p-value 3.6982e-13
- # Data 183
- Norm. par. 17
- Int. Method KEPLER
- Min. Method SIMPLEX
- Element type ASTROCENTR

```

=> kremove.data(k, 'all'); k$star <- 0.950000; k$epoch
<- NaN; kadd.data(k, c("./datafiles/HD115617_AA.vels",
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[1] 10.60109
=> b_1 <- kmcmc(k, chains=2, R.stop=1.1, skip.
first=1000, discard=k$nrpars*10, print=TRUE)
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Rsingle_max = 1.488546e-01
# mcmc, chains = 2, R.stop = 1.100000e+00, start =
perturb, noise=TRUE, skip = 1000, discard = 170, tot.
length = 10472

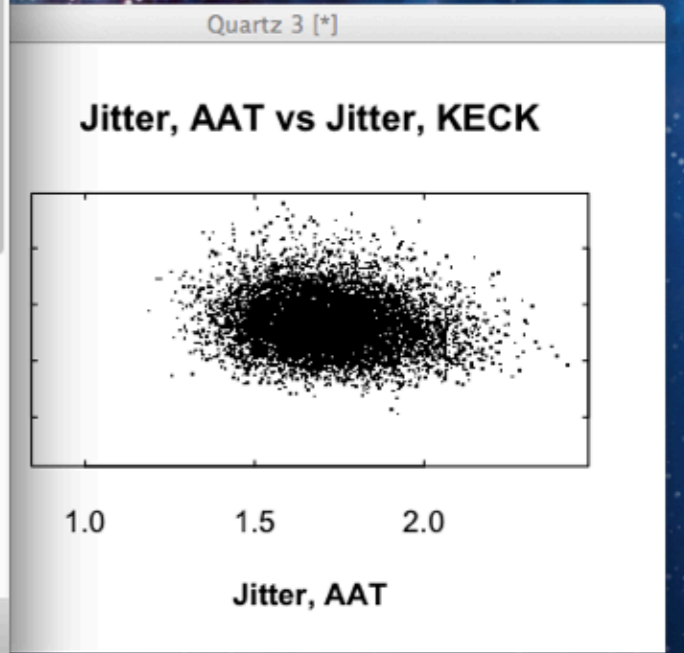
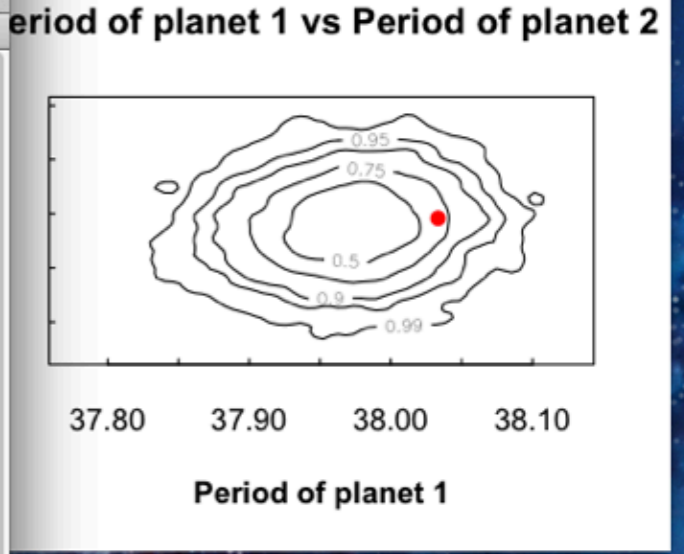
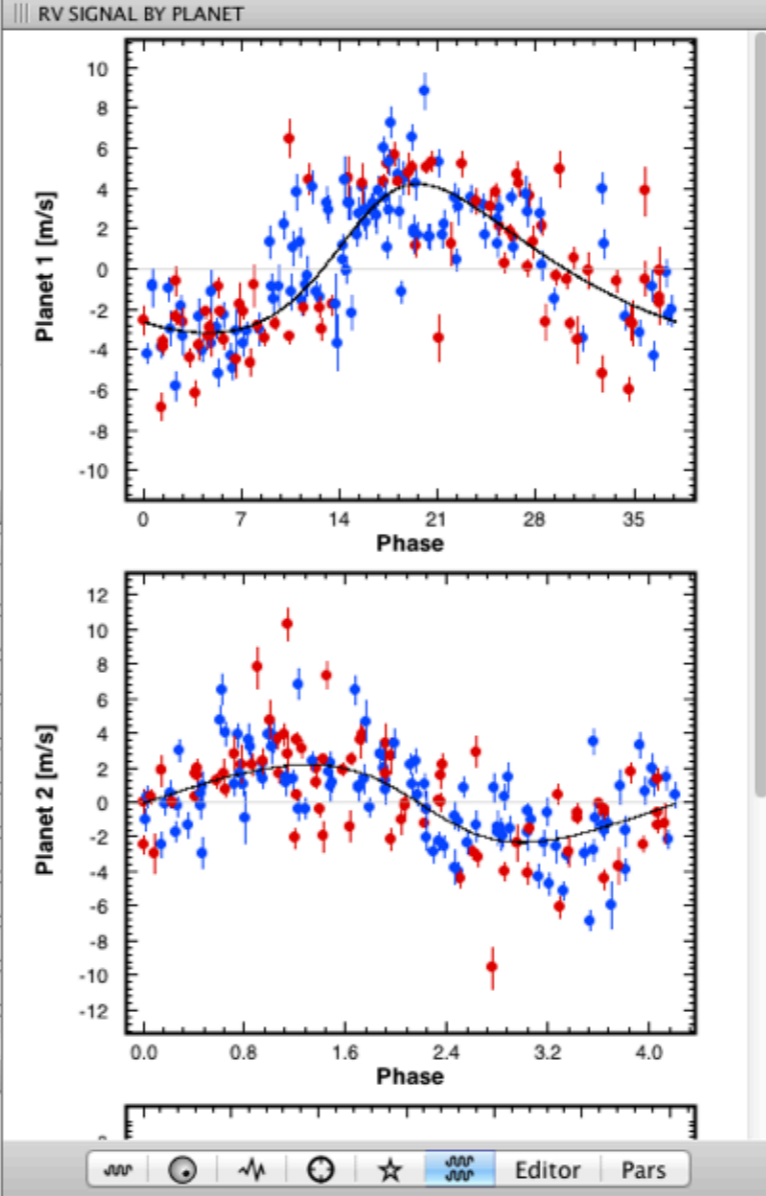
plot(b_1, type='scatter', px=list('par','data.noise1'), p
y=list('par','data.noise2'), dev.factor=5, pch='.', xlab=
'Jitter, AAT', ylab='Jitter, KECK')
                    
```

ORBITAL ELEMENTS

	Planet 1	Planet 2	Planet 3
Period [d]	38.03334	4.21492	121.55729
Mass [Mj]	0.05754	0.01702	0.06255
Mean anomaly [deg]	206.54303	166.00596	40.56638
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Inclination [deg]	90	90	90
Node [deg]	0	0	0
Semi-maj. axis [AU]	0.21759	0.0502	0.47212
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Radius [Rj]	0	0	0

DATA PARAMETERS

HD115617_... 0.2344 HD115617_... 0.9892



Model statistics (Chi², log likelihood, etc.)

Command line

Plots (interactively updated)

PROPERTIES

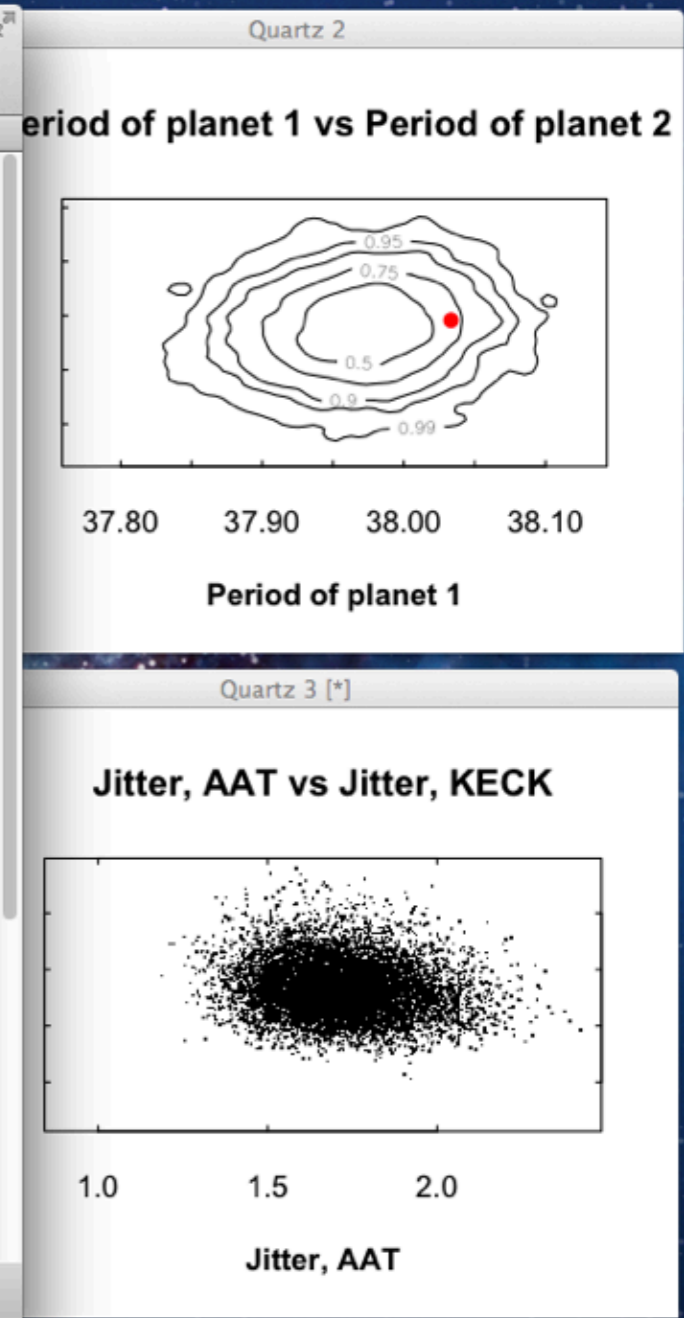
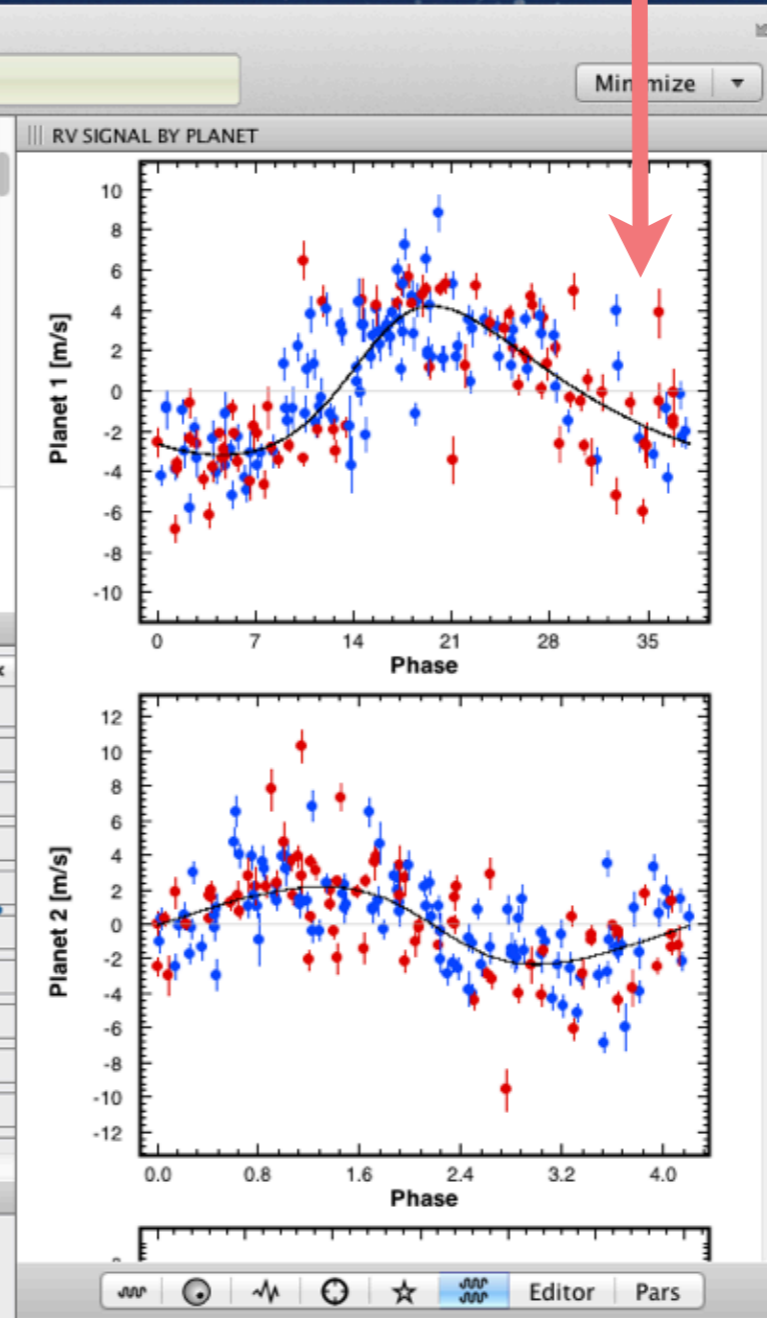
- # Planets: 3
- Star mass [M_⊙]: 0.95
- Epoch [JD]: 2453369.16
- Chi²_{RV}: 1759.78016
- RMS_{RV} [m/s]: 2.16691
- Jitter [m/s]: 2.04813
- # RVs: 183
- Chi²_{TT}: 0
- RMS_{TT} [d]: --
- # TTs: 0

ORBITAL ELEMENTS

	Planet 1	Planet 2	Planet 3
Period [d]	38.03334	4.21492	121.55729
Mass [M _J]	0.05754	0.01702	0.06255
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Inclination [deg]	90	90	90
Node [deg]	0	0	0
Semi-maj. axis [AU]	0.21759	0.0502	0.47212
Semi-amplitude [m/s]	3.68719	2.24459	2.83278
Radius [R _J]	0	0	0

DATA PARAMETERS

	0.2344	0.9892
HD115617_...		
HD115617_...		



Model statistics (Chi², log likelihood, etc.)

Command line

Plots (interactively updated)

PROPERTIES

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- Star mass [M_⊙]: 0.95
- Epoch [JD]: 2453369.16
- Chi²_{RV}: 1759.78016
- RMS_{RV} [m/s]: 2.16691
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- # RVs: 183
- Chi²_{TT}: 0
- RMS_{TT} [d]: --
- # TTs: 0

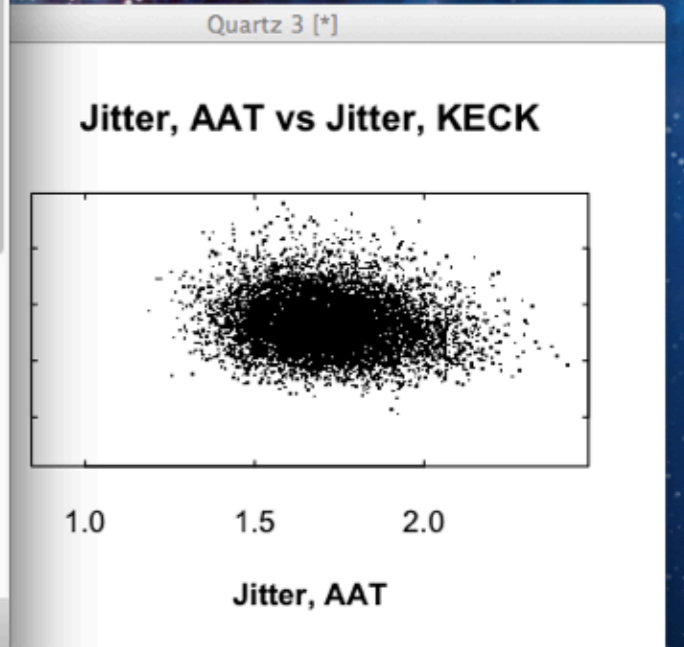
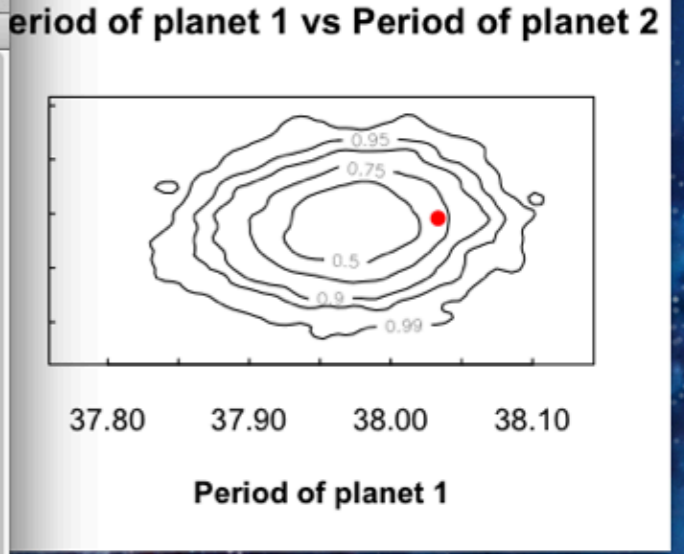
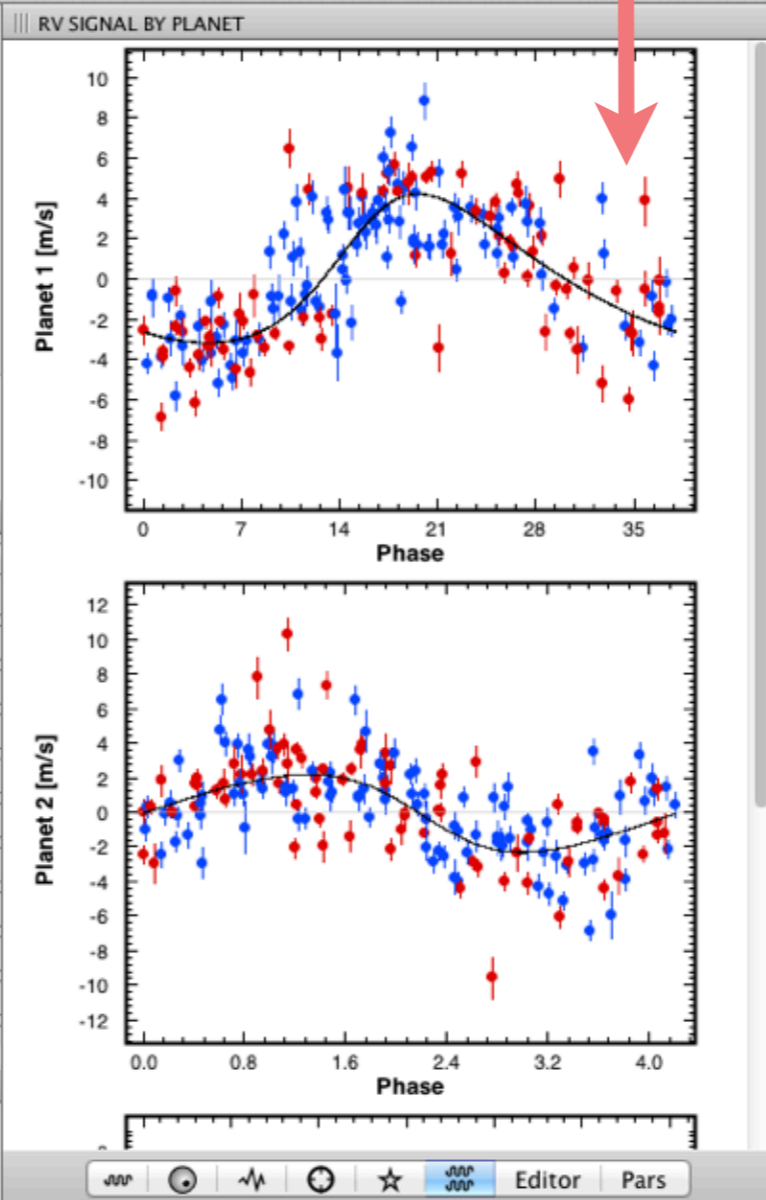
ORBITAL ELEMENTS

	Planet 1	Planet 2	Planet 3
Period [d]	38.03334	4.21492	121.55729
Mass [M _J]	0.05754	0.01702	0.06255
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Eccentricity	0.21819	0.16019	0.34807
Long. Peri [deg]	311.01757	100.3883	311.42328
Inclination [deg]	90	90	90
Node [deg]	0	0	0
Semi-maj. axis [AU]	0.21759	0.0502	0.47212
Semi-amplitude [m/s]	3.68719	2.24459	2.83278
Radius [R _J]	0	0	0

DATA PARAMETERS

	0.2344	0.9892
HD115617_...		

```
=> kremove.data(k, 'all'); k$star <- 0.950000; $epoch <- NaN; kadd.data(k, c("./datafiles/HD115617_AA.vels", "./datafiles/HD115617_KECK.vels")); => kminimize(k) [1] 10.60109 => b_1 <- kmcmc(k, chains=2, R.stop=1.1, skip.first=1000, discard=k$nrpars*10, print=TRUE) Final length: 5236, final R_max = 1.065146e+00, final Rsingle_max = 1.488546e-01 # mcmc, chains = 2, R.stop = 1.100000e+00, start = perturb, noise=TRUE, skip = 1000, discard = 170, tot.length = 10472 plot(b_1, type='scatter', px=list('par','data.noise1'), py=list('par','data.noise2'), dev.factor=5, pch='.', xlab='Jitter, AAT', ylab='Jitter, KECK')
```



Orbital plot



Model statistics (Chi², log likelihood, etc.)

Command line

Plots (interactively updated)

The screenshot displays a software interface for exoplanet modeling. It is divided into several sections:

- Model statistics:** Located on the left, it lists various statistical values such as χ^2_{RV} , RMS_{RV}, Jitter, χ^2_{TT} , RMS_{TT}, χ^2 , $\chi^2_{reduced}$, -Log likelihood, KS p-value, # Data, Norm. par., Int. Method, Min. Method, and Element type.
- Command line:** The central area shows R code for data loading, model fitting, and plotting. A red arrow points to the `kminimize(k)` command.
- Orbital Elements:** A table below the command line lists parameters for three planets (Planet 1, Planet 2, Planet 3), including Period [d], Mass [Mj], Mean anomaly [deg], Eccentricity, Long. Peri [deg], Inclination [deg], Node [deg], Semi-maj. axis [AU], Semi-amplitude [m/s], and Radius [Rj]. A red arrow points to the Radius [Rj] column.
- Plots:** On the right, there are three plots:
 - RV SIGNAL BY PLANET:** Two scatter plots showing radial velocity signals for Planet 1 and Planet 2. A red arrow points to the top plot.
 - Contour Plot:** A plot titled "period of planet 1 vs Period of planet 2" showing contours of probability density. A red dot indicates the best-fit parameters.
 - Scatter Plot:** A plot titled "Jitter, AAT vs Jitter, KECK" showing the relationship between jitter measurements from two different telescopes.
- Orbital plot:** A small circular icon in the bottom left corner, with a red arrow pointing to it.

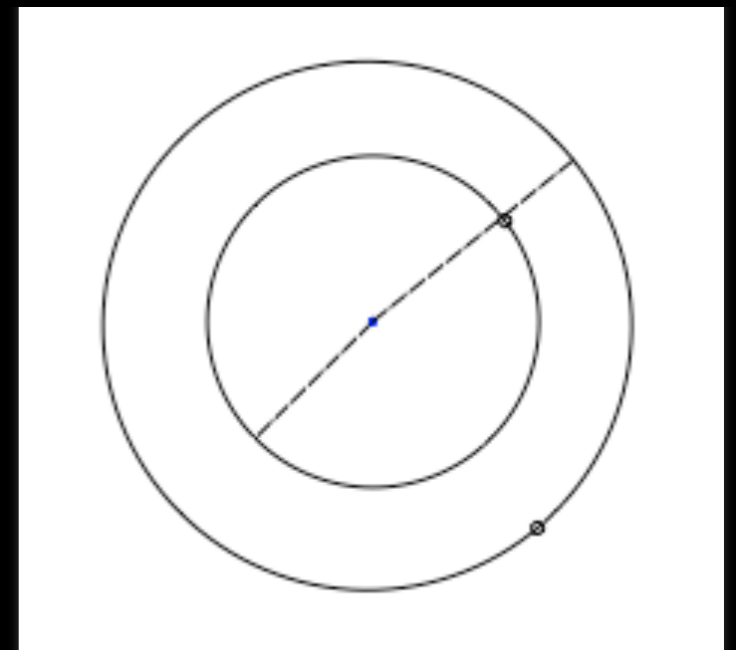
Orbital plot

Model parameters

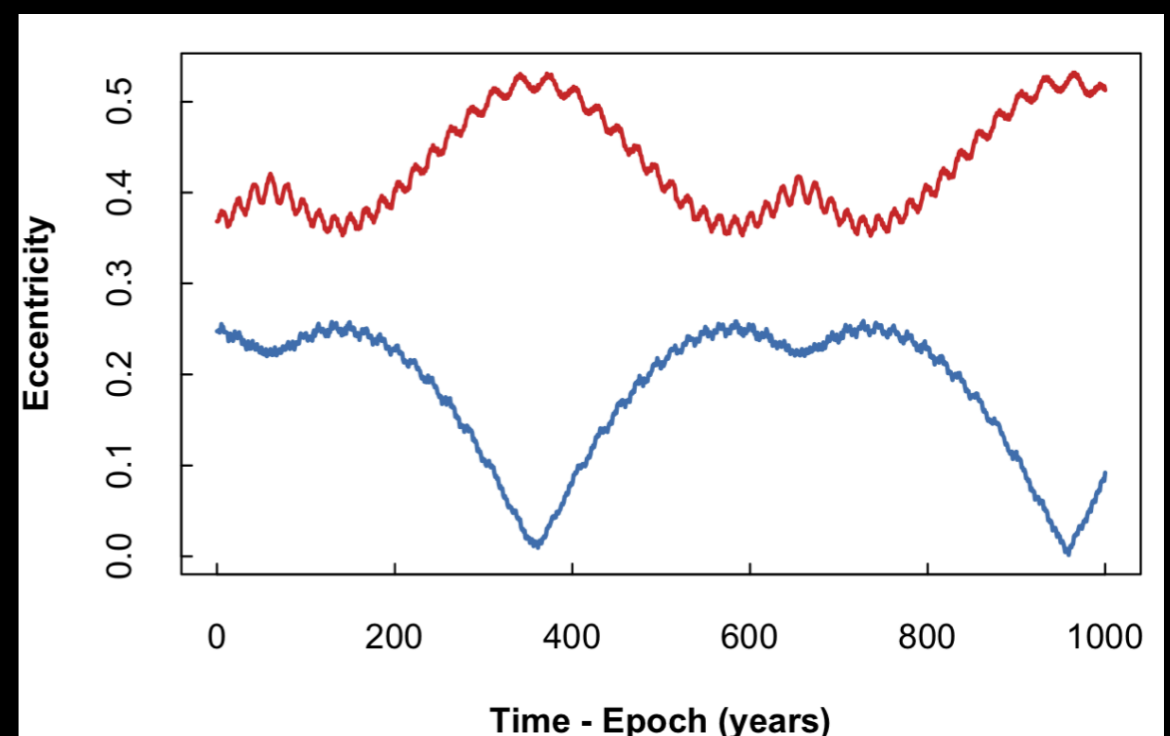
Dynamical fitting

Models can optionally include gravitational interactions between bodies:

(1) Fit strongly interacting/resonant systems (e.g. GJ876, HD128311, etc.)

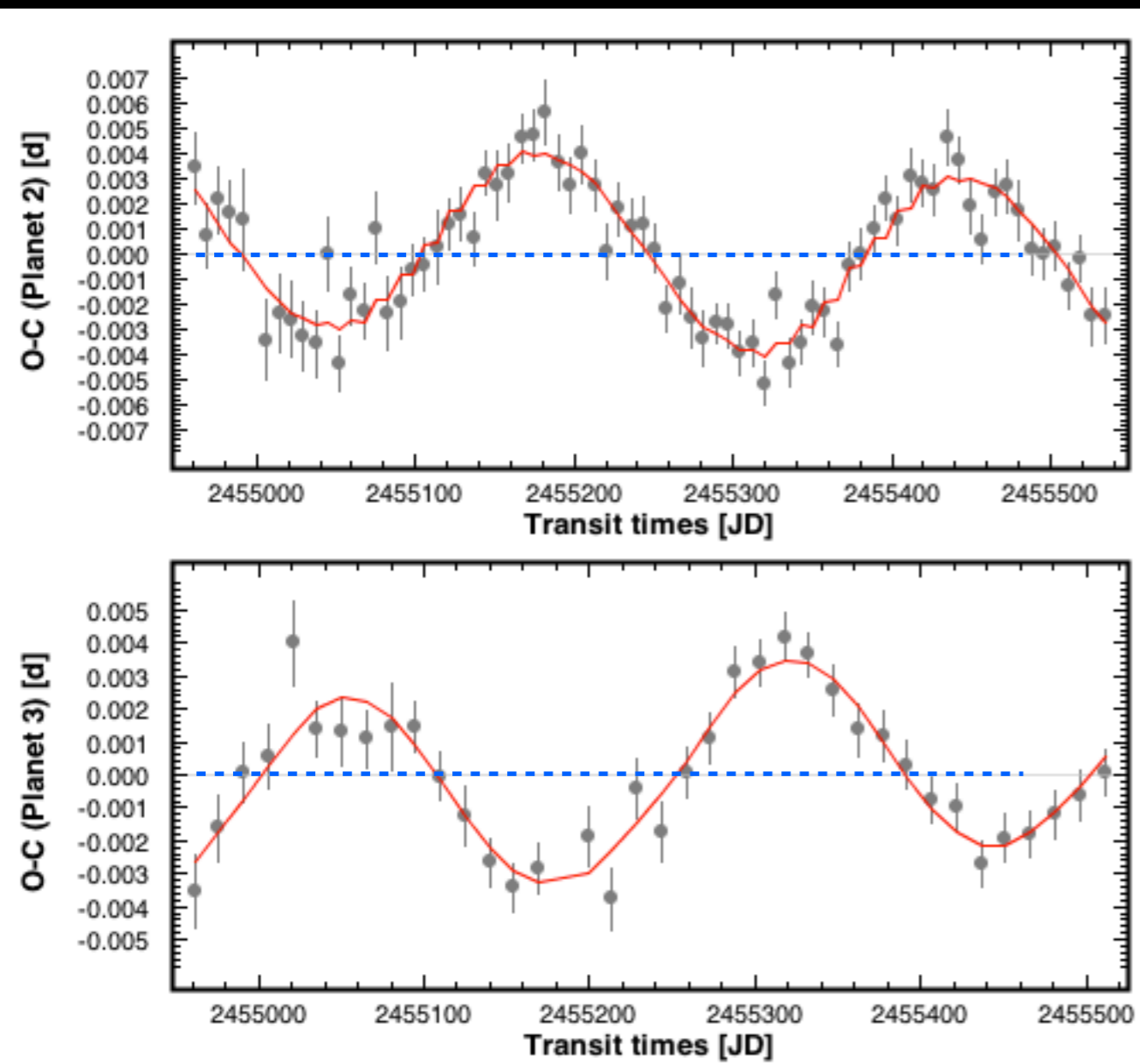


(2) Check for the long-term stability of a planetary system and create stability maps

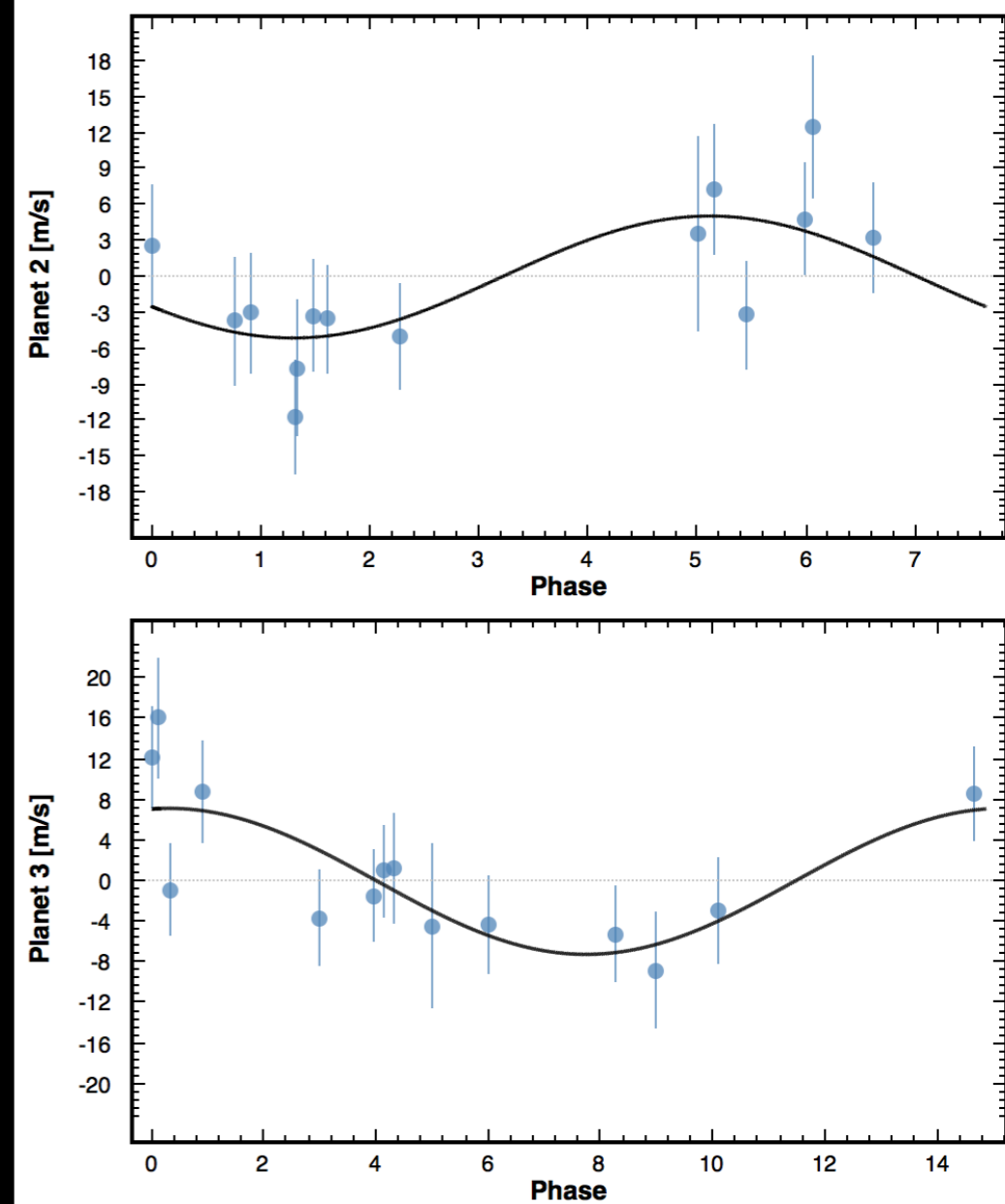


Dynamical fitting

(3) Fit transit timing datasets combined with radial velocity datasets and take advantage of **transit timing variations** to constrain orbital elements.



+



- Lomb-Scargle and bootstrapped periodogram
- Keplerian and self-consistent fitting
- Long-term integration using SWIFT
- Optimization using Simplex, Levenberg-Marquardt, Simulated Annealing or Differential Evolution
- Error estimation using Markov-Chain Monte Carlo or bootstrap
- Model cross-validation using jack-knife
- Completely customizable models (e.g., add new parameters to the model)
- Algorithms are automatically parallelized to run across multiple cores; some algorithms can run across computing clusters
- And more!

Demo

It'll be quick, I promise!

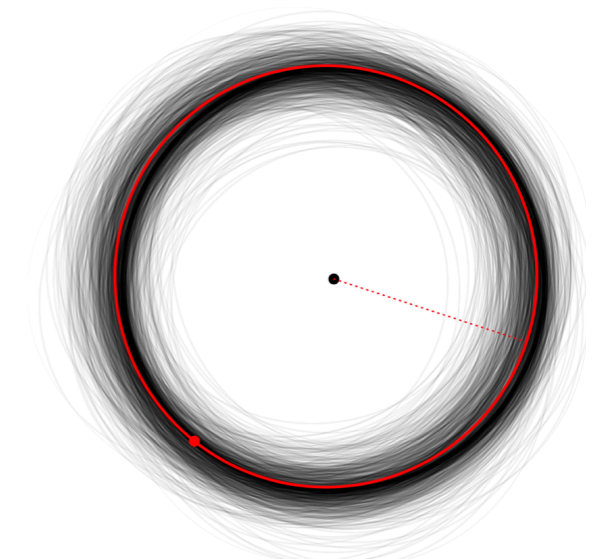
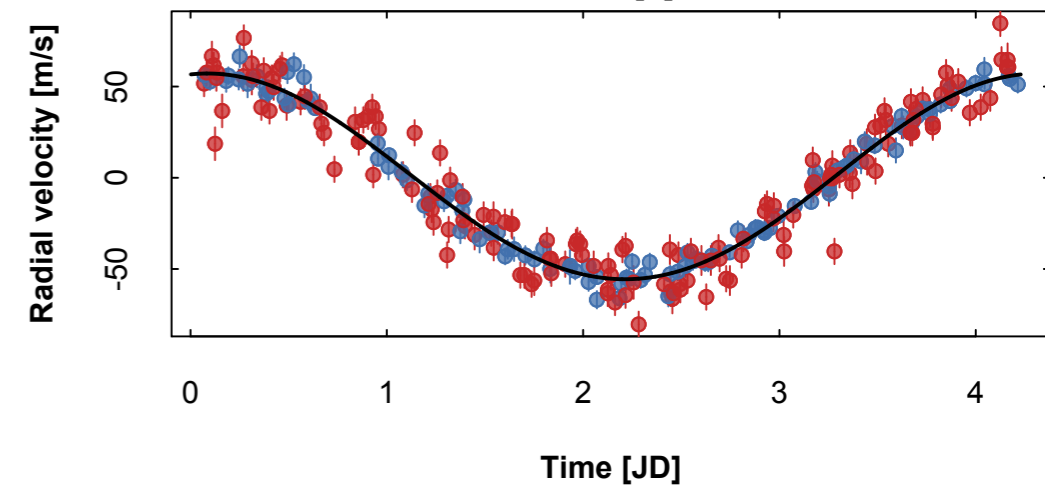
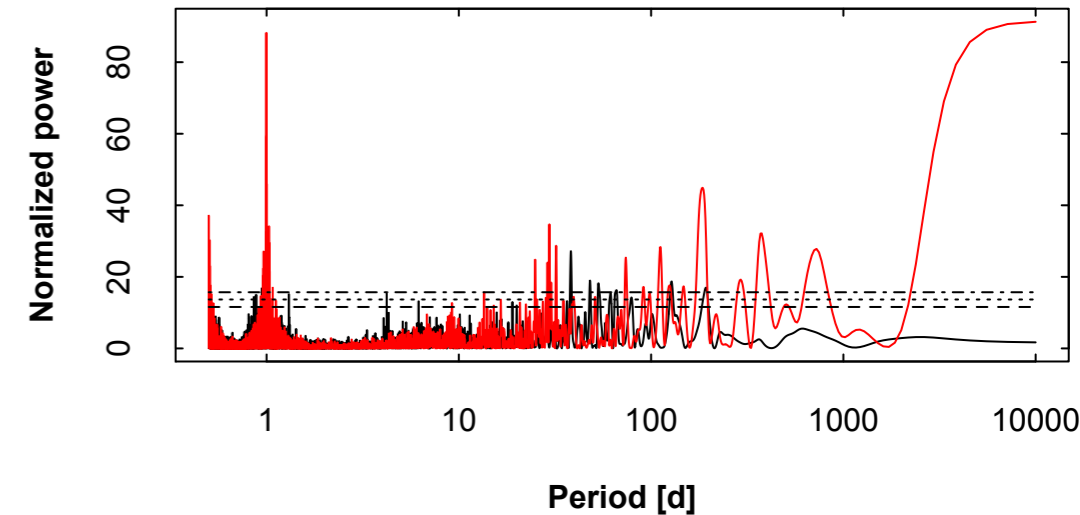
Systemic is also an R package.

This means that you can write full-fledged scripts to analyze your data, and interface with literally thousands of sophisticated statistical packages.

Computations are parallelized and can be run across clusters.

Load data, add a planet and run a Markov-Chain Monte Carlo algorithm.

```
# Creates a new model object  
k <- knew()  
# Load new data  
kadd.data(k, "1p1.vels")  
# Calculate the power spectrum of the data  
p <- kperiodogram(k)  
  
# Add a planet at the period corresponding  
# to the highest peak  
kadd.planet(k, c(period = p[1, 'period']))  
kminimize(k)  
plot(k)  
  
# Run a Markov-Chain Monte Carlo analysis  
# (with default parameters)  
kmcmc(k, chains=5)
```



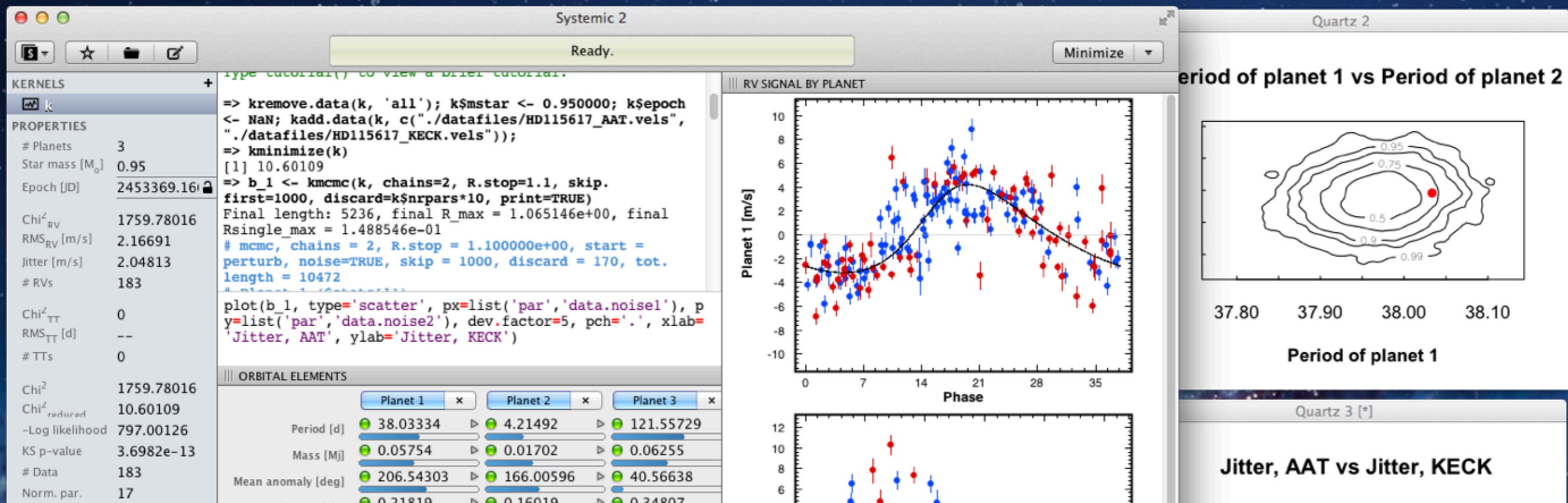
Systemic 2

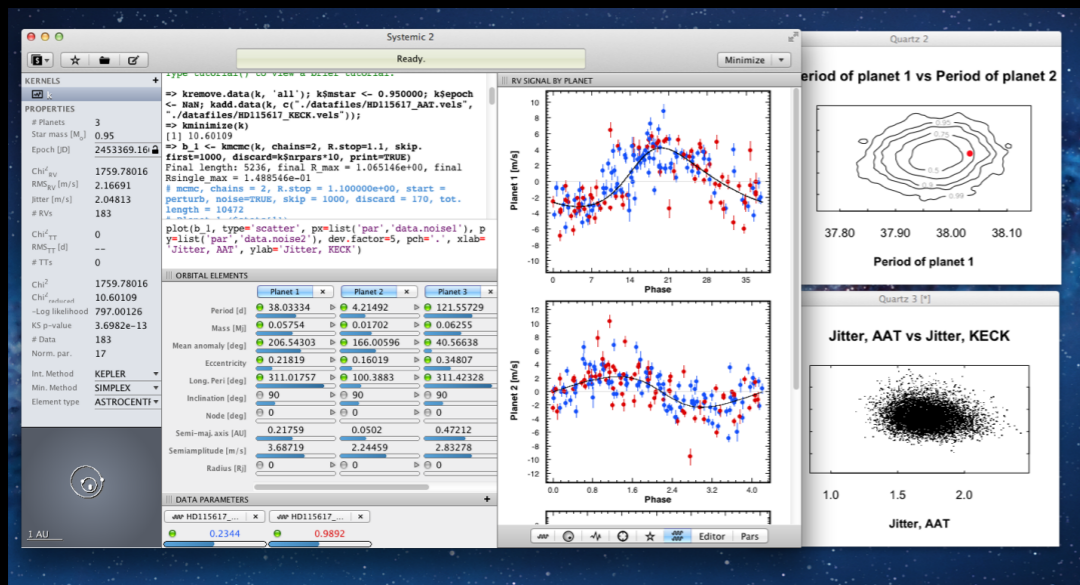
This package and its source code is free and available on GitHub: anyone can download it and modify it freely...

<http://github.com/stefano-meschiari/Systemic2>

Teaching & Outreach

One could use the “full” Systemic to let students analyze exoplanetary data, but its interface can be overwhelming...





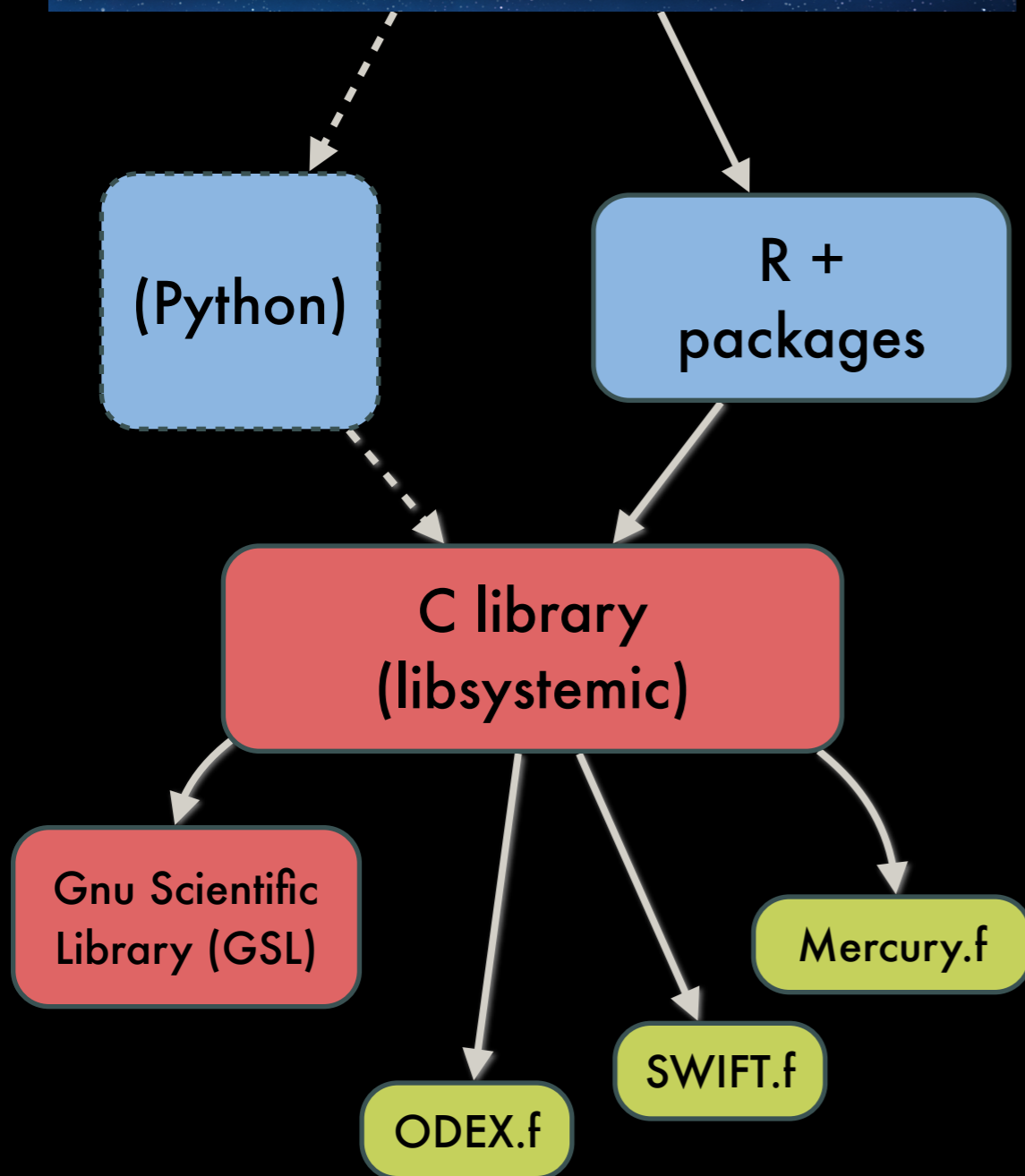
User interface (Java)

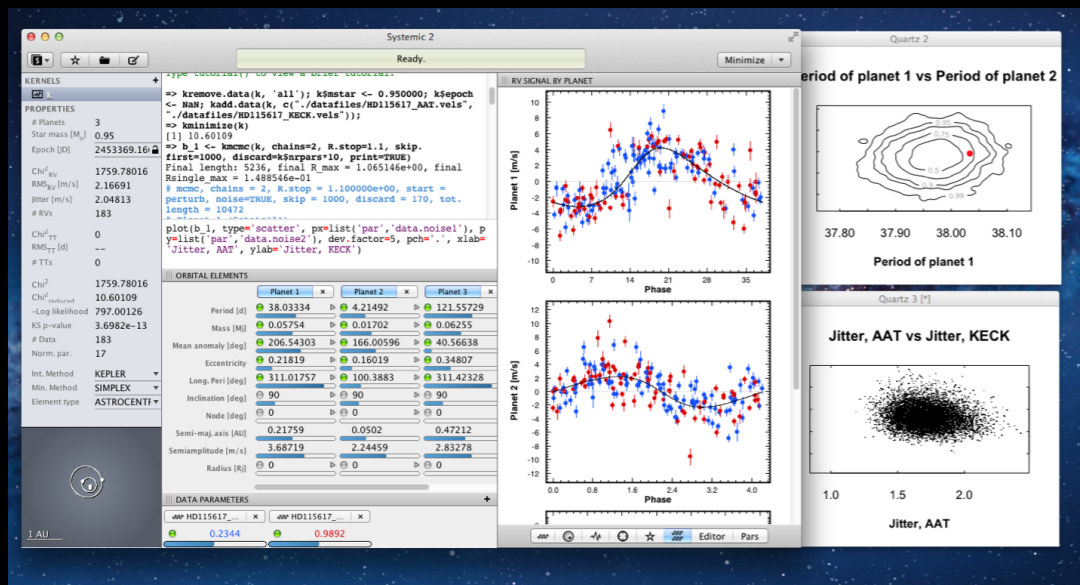
High-level language (R and packages)

Core code:

C library

+ some Fortran code





User interface (Java)

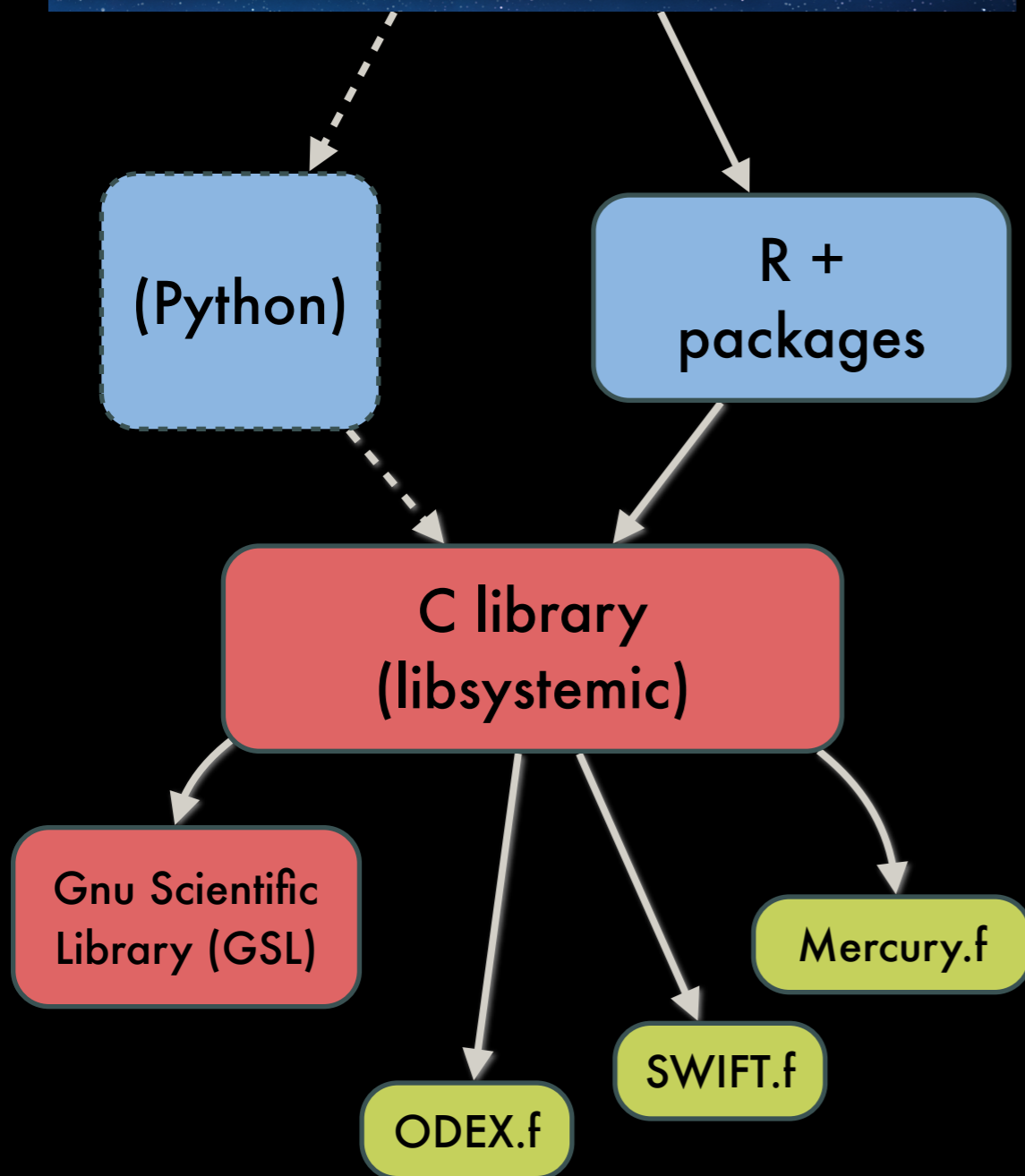
High-level language (R and packages)

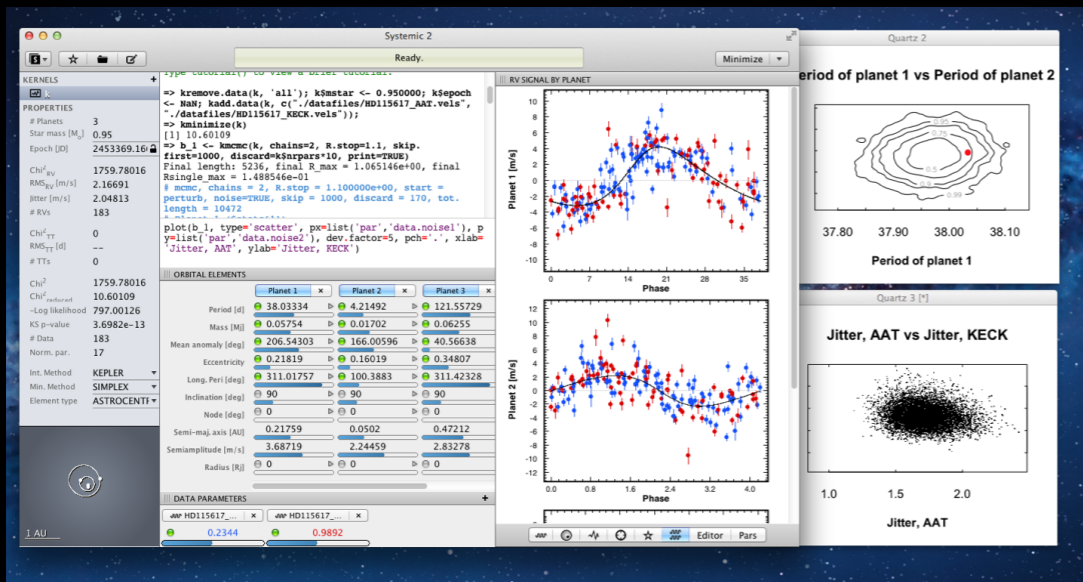
Core code:

C library

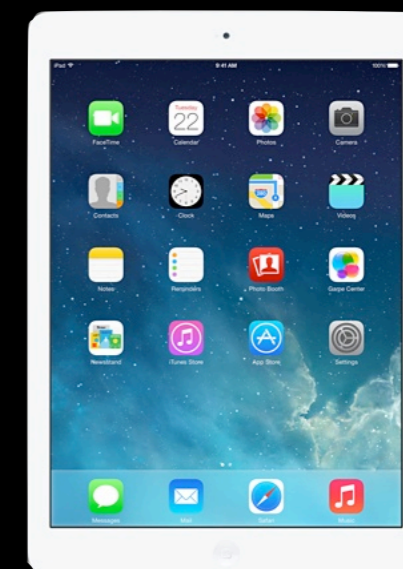
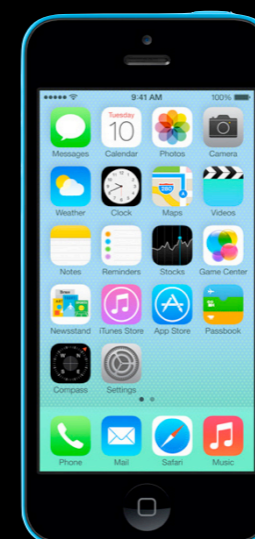
+ some Fortran code

Best installation experience is no installation.





The Web



(Python)

R + packages

C library (libsystemic)

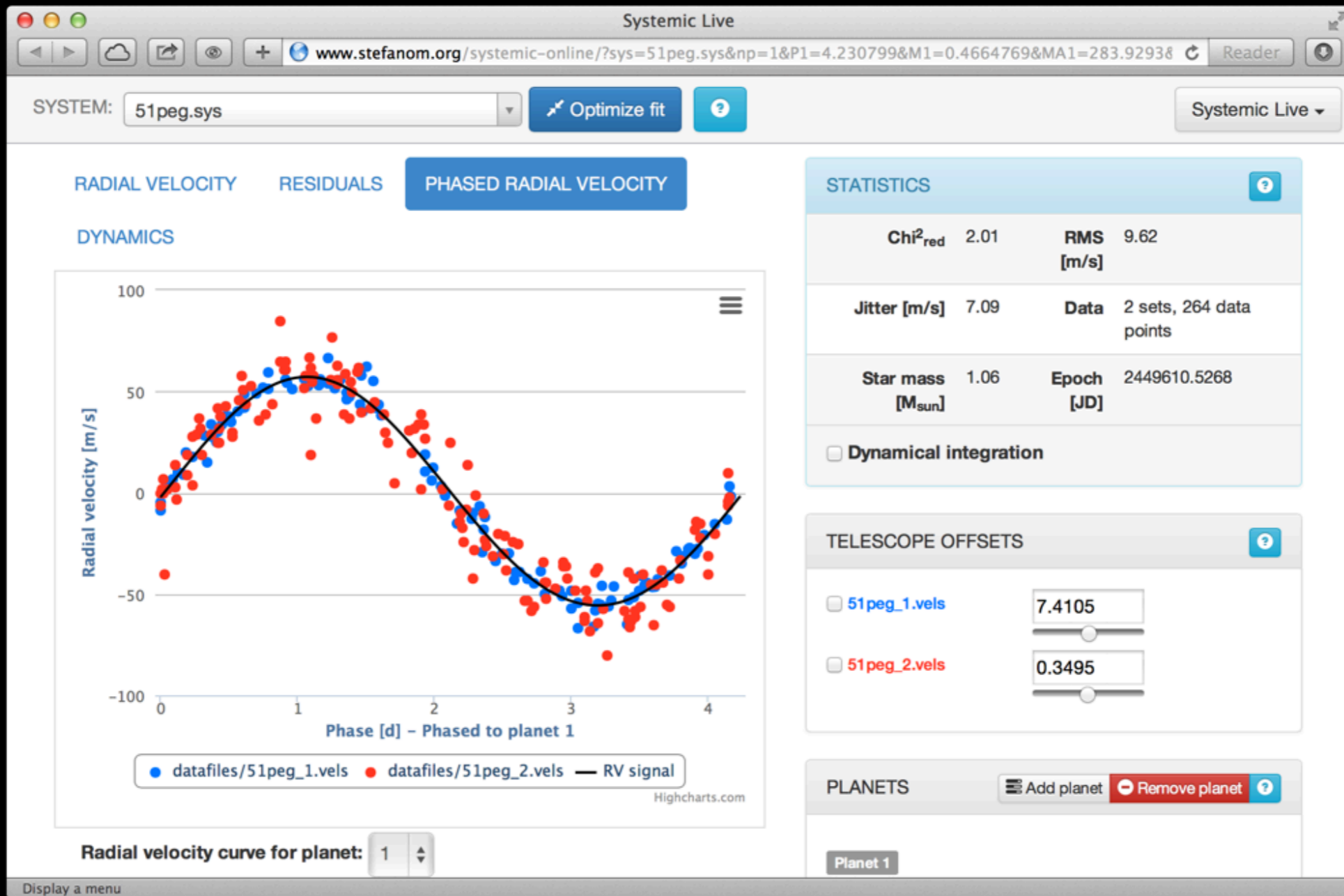
Gnu Scientific Library (GSL)

Mercury.f

ODEX.f

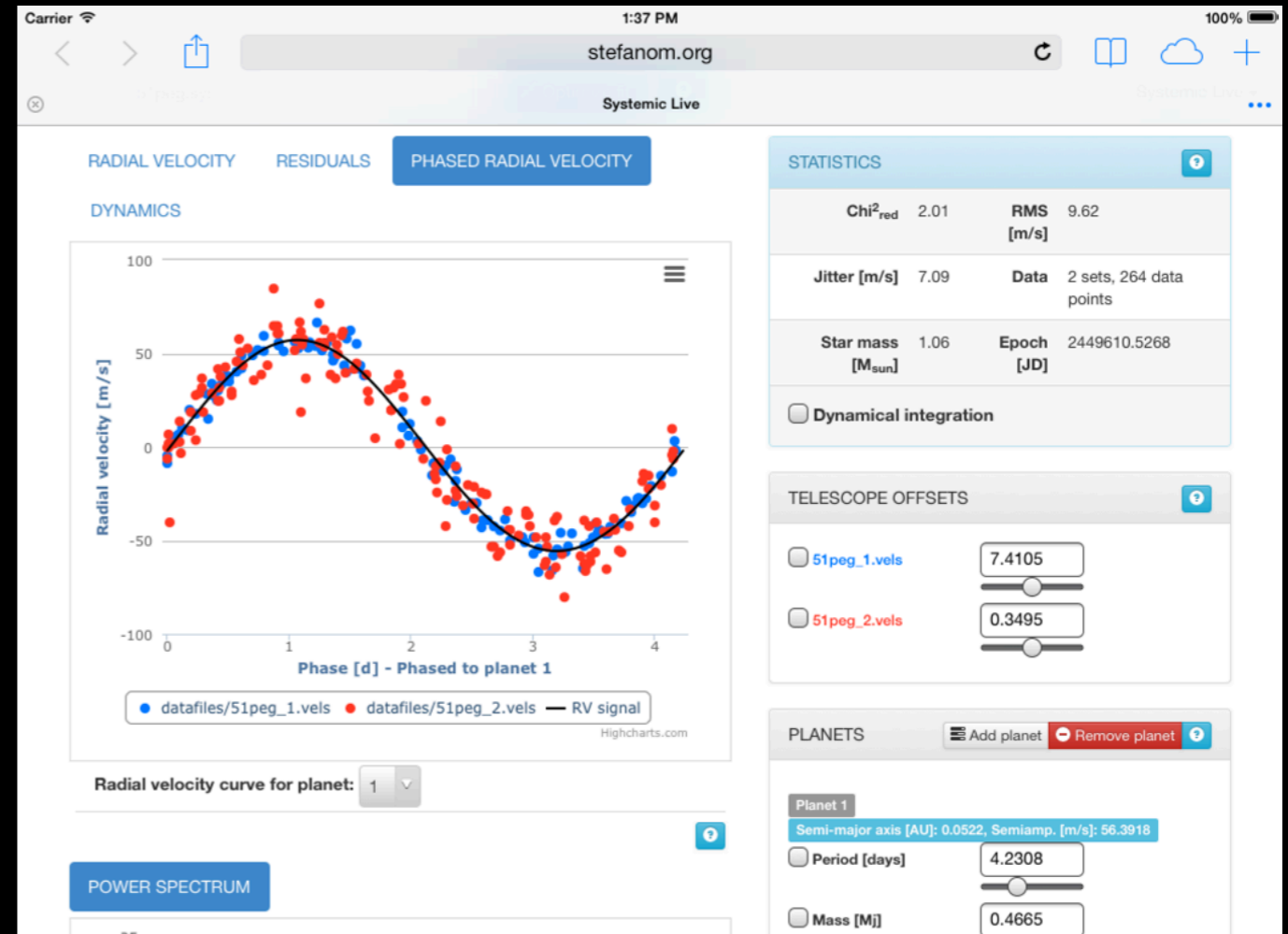
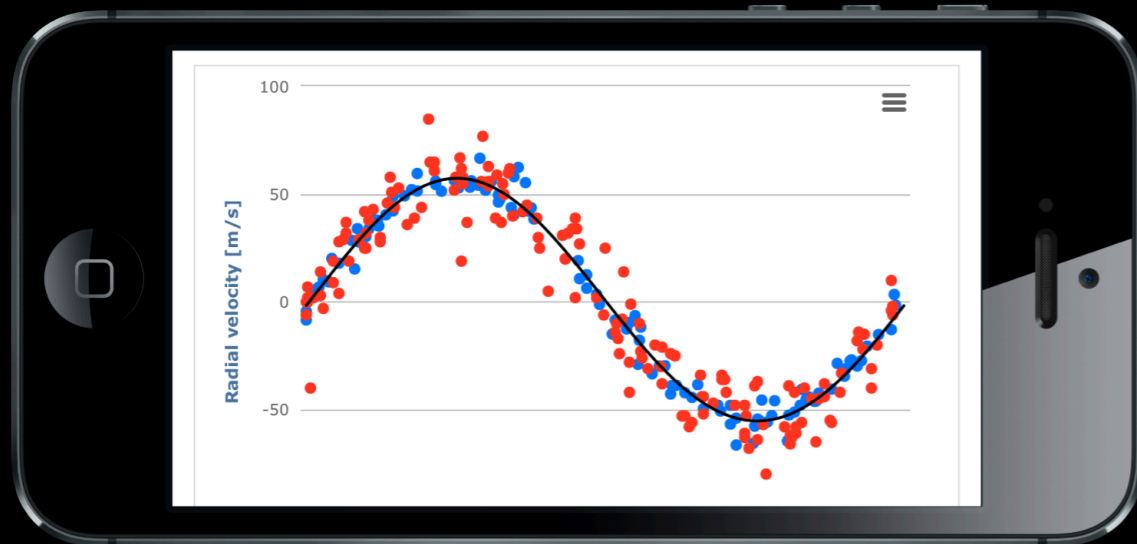
SWIFT.f

Systemic Live



A simplified web app for modelling exoplanetary data, at the just the right level for high school & undergraduate classes.

Systemic Live



Works on smartphones and tablets, too.

Demo

<http://www.stefanom.org/systemic-online/?sys=51peg.sys>

Systemic Live

It's easy to share a planetary model using just the current web address, like so:

[http://www.stefanom.org/systemic-online/?
sys=51peg.sys&np=1&P1=4.230799&M1=0.466
4769&MA1=283.9293&E1=0.0140892&L1=34
4.9533&o0=7.410511&o1=0.3495096&im=0](http://www.stefanom.org/systemic-online/?sys=51peg.sys&np=1&P1=4.230799&M1=0.4664769&MA1=283.9293&E1=0.0140892&L1=344.9533&o0=7.410511&o1=0.3495096&im=0)

“The online Systemic Console is a real gift to the community. The online console distills years of work to optimize the modeling real radial velocity data. Students can run bootstrap Monte Carlo codes to determine measurement errors and numerical integrations to determine the dynamical stability of multi-planet systems. I use this site to train both undergraduate and graduate students – they love the power of this program.”

Debra Fischer, Yale University

“Systemic is simple enough to use that it can provide a hand-on ‘virtual lab’ for a large general education class, [...] students can get a taste of the scientific process even before they learn to program” –

Eric Ford, Penn State

“I have used Systemic for several years in my class for advanced undergraduate physics majors. The students favorite problem set uses Systemic to explore real radial velocity data sets and compare their solutions to orbital parameters for published systems. Systemic is extremely sophisticated, but easy to use, so it allows students to get a feeling for the tools used in real exoplanet research.”

Jonathan Fortney, UC Santa Cruz

Tutorials & labs

You can find tutorials and labs on the webpage:

<http://www.stefanom.org/systemic-live/>

- **“51 Pegged: Rediscovering the First Exoplanet”**
- **“A Fish in a Barrel – HD 4208b”**
- **“The Ups and Downs of Ups And”**
- **...and others**

I'm also working to connect existing exoplanet “databases” (exoplanet.eu, exoplanet.org, etc.) with Systemic, so that with a click you can access and analyze the data associated with a system.

HOWTO

My biased recommendations on getting it done, as a busy astronomer who's eager to learn new valuable skills and do some outreach in the process.

<https://github.com/stefano-meschiari/Notes>

Making online web apps

Presentational part: Structure, Layout and Appearance



Making online web apps

Presentational part:

- **Learn the very basics of HTML5 and CSS.**

You should start grokking this stuff anyway if you're making your own webpage (see Chalence's talk). HTML defines the structure of the page, CSS its appearance (roughly).

Recommendation: the **Mozilla Developer Network (MDN)** is really great place to start. Unfortunately, a lot of bad/misleading resources bubble up Google searches (*cough* W3Schools *cough*), so beware.

Making online web apps

Presentational part:

- **Make a framework do most of the work.**

A framework takes care of a lot of things that are objectively complicated even in modern browsers, like complex layouts, components, smoothing over browser differences, and more. They are usually pretty quick to learn and give your projects a professional look.

Recommendation: **Bootstrap** or **UIKit**.

Making online web apps

Interactivity/computations

```
systemic.js (/Users/sm52286/Projects/SystemicLive/js/systemic.js)
944
945   var optimize = function() {
946     var active = false;
947     for (var i = 1; i <= K_getNplanets(k); i++) {
948       for (var j = PER; j <= LOP; j++) {
949         if ($("#elementSel_" + i + "_" + j).is(":checked")) {
950           K_setElementFlag(k, i, j, MINIMIZE+ACTIVE);
951           active = true;
952         }
953         else
954           K_setElementFlag(k, i, j, ACTIVE);
955       }
956     }
957     for (var i = 0; i < MAX_SETS; i++)
958       if ($("#offsetSel_" + i).is(":checked")) {
959         K_setParFlag(k, i, MINIMIZE+ACTIVE);
960         active = true;
961       }
962     else
963       K_setParFlag(k, i, ACTIVE);
964
965     if (!active) {
966       alert("You need to have at least one parameter selected next to each parameter.");
967       return;
968     }
969
970
971     var chi2 = K_getChi2(k);
```

 Optimize fit

Making online web apps

Interactivity/computations:

- **Learn JavaScript.**

JavaScript superficially looks very similar to C or Java.

```
function square(x) {  
    return x * x;  
}
```

```
var x = 2;  
alert("The square of" + x + " is " + sqr(x));
```

In reality, very different conceptually and functionally.

Making online web apps

Interactivity/computations:

- **Learn JavaScript.**

In many ways, it's an evil, evil language. At the same time, *only* language allowed on browsers (no C/Fortran/IDL/Python/Perl/Ruby/anything), so it's an incredibly valuable skill.

This is probably the **hardest component to learn correctly.**

Making online web apps

Interactivity/computations:

Recommendations:

Making online web apps

Interactivity/computations:

Recommendations:

- **JavaScript: the Good Parts** is brief, clear and warns you about warts & pitfalls of the language.

Making online web apps

Interactivity/computations:

Recommendations:

- **JavaScript: the Good Parts** is brief, clear and warns you about warts & pitfalls of the language.
- **Mozilla Developer Network** is a useful reference for JavaScript as well.

Making online web apps

Interactivity/computations:

Recommendations:

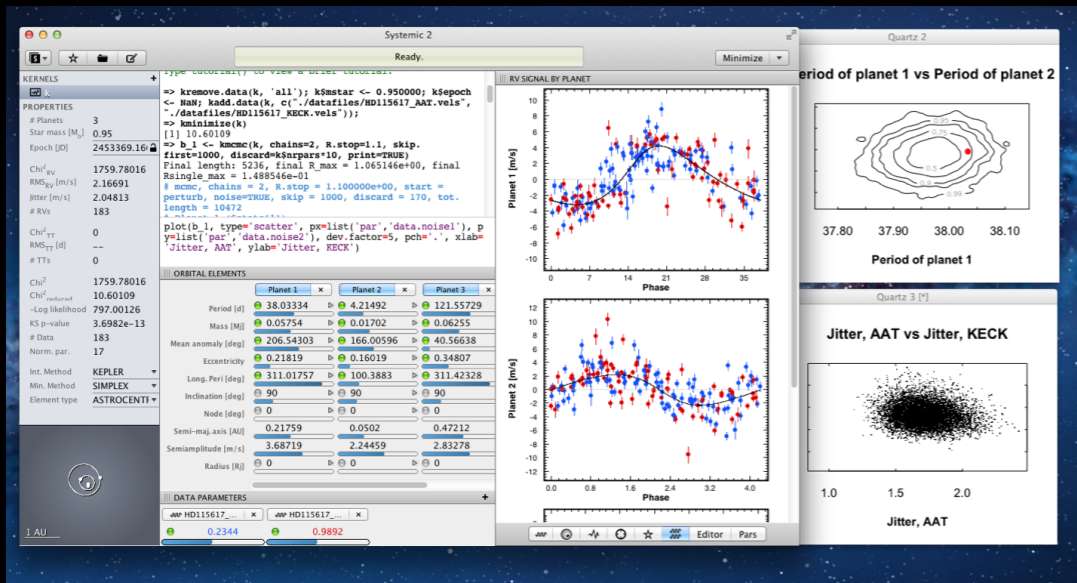
- **JavaScript: the Good Parts** is brief, clear and warns you about warts & pitfalls of the language.
- **Mozilla Developer Network** is a useful reference for JavaScript as well.
- Again, frameworks and libraries can lessen the pain and make you more productive. **jQuery** and **Underscore.js** have taken a lot of the friction out of interacting with the webpage elements for me.

Making online web apps

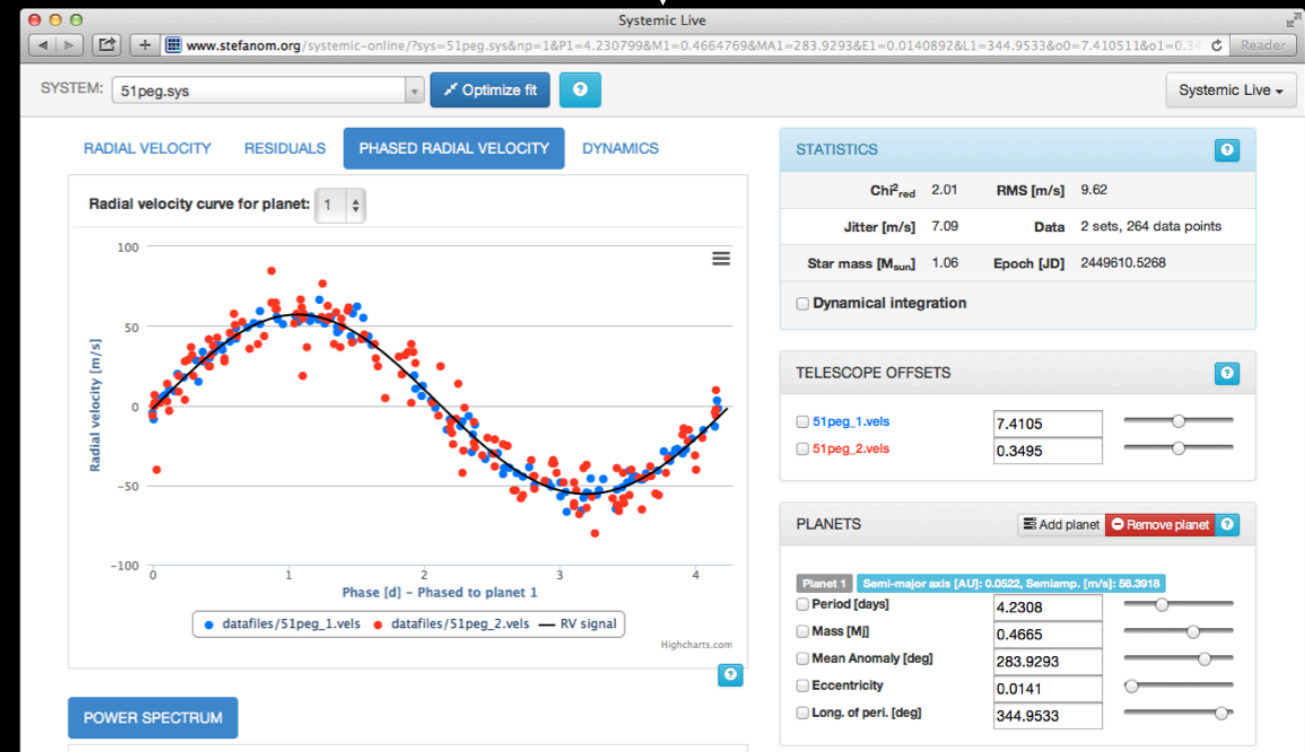
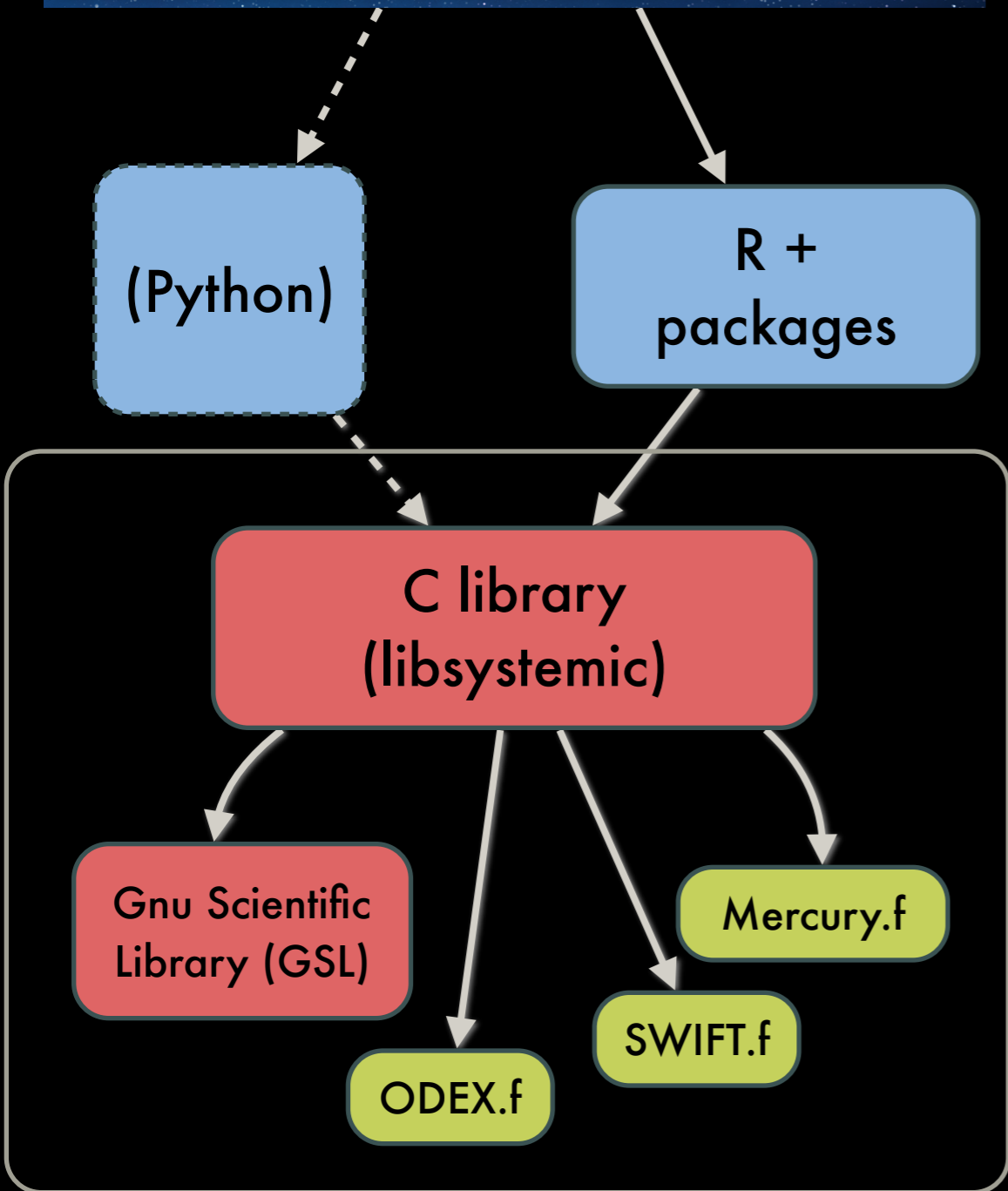
Interactivity/computations:

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- **JavaScript: the Good Parts** is brief, clear and warns you about warts & pitfalls of the language.
- **Mozilla Developer Network** is a useful reference for JavaScript as well.
- Again, frameworks and libraries can lessen the pain and make you more productive. **jQuery** and **Underscore.js** have taken a lot of the friction out of interacting with the webpage elements for me.
- **Node.js** is a way to run and test your JavaScript *outside* your browser.



emscripten
converts C code into Javascript



Making online web apps

Interactivity/computations:

Recommendations, part deux:

- **Emscripten** is a fabulous way to translate complex but trusty C (Fortran) code into Javascript code. You literally could just change this command:

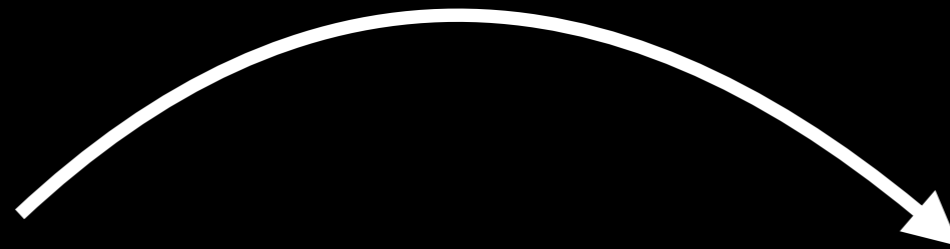
```
make MyProject
```

into this:

```
emmake MyProject
```

and you would get Javascript (instead of machine code) as the output.

Systemic Live



Science

Teaching &
outreach

Any improvement to the scientific software goes directly into the teaching & outreach code.

Making online web apps

Plotting:

- **Lots of very different approaches.**

I tend to prefer more limited libraries that are “turn-key”, i.e. do not require to learn a whole different paradigm just for plotting some data. E.g. just specify that you want a scatterplot, provide the data and go.

Recommendation: **Highcharts** (free for edu)

**One more outreach
thingy**

How do you reach people that are very enthusiastic about exoplanets, but don't have the technical skills, or patience, or interest in looking at real data?

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We have all this really good code and a way to port it on the Web, so what do you do?

How do you reach people that are very enthusiastic about exoplanets, but don't have the technical skills, or patience, or interest in looking at real data?

We have all this really good code and a way to port it on the Web, so what do you do?

Make a game!

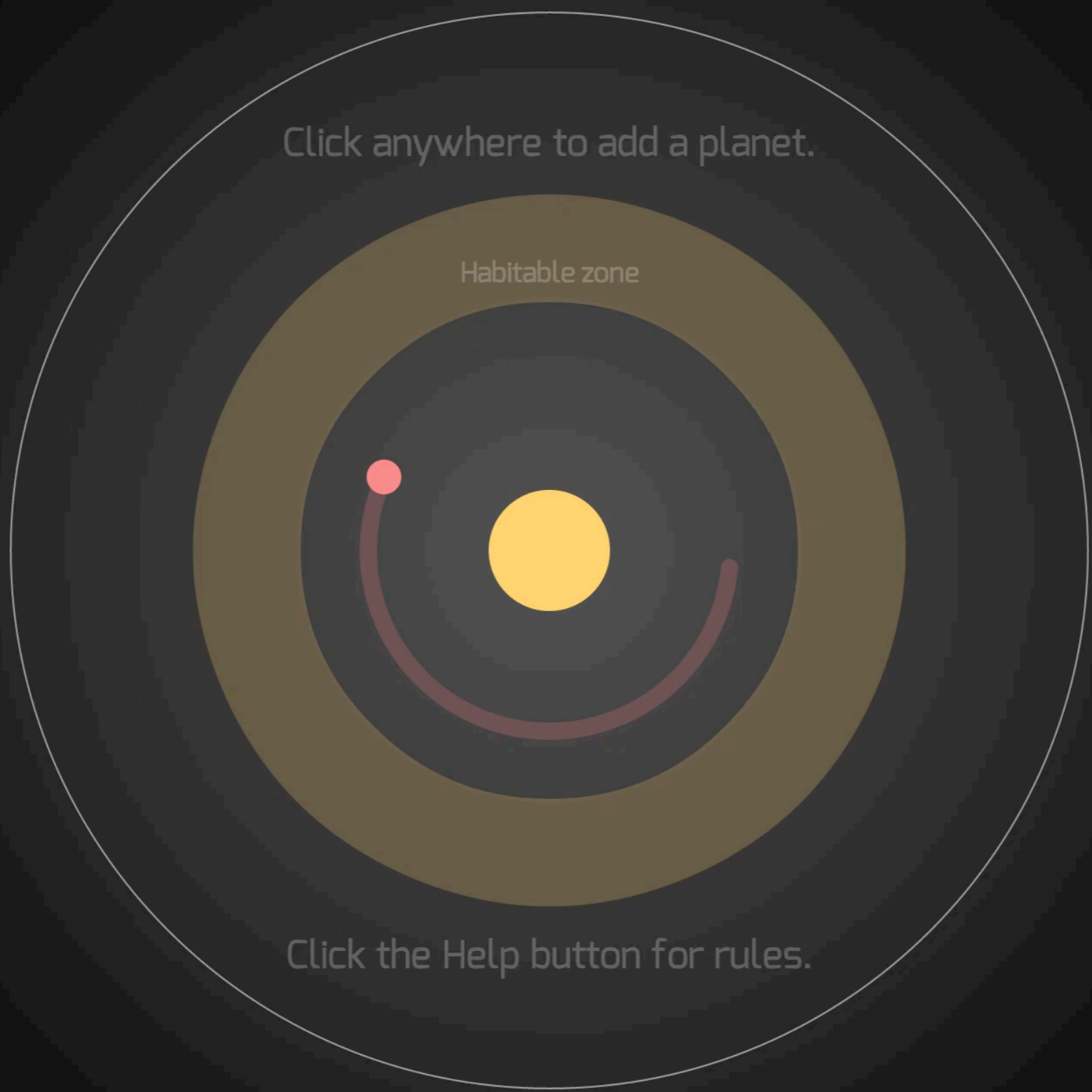
Super Planet Crush

Click on the type of body to add next:

- Earth **1x**
- Super-Earth **5x**
- Ice giant **15x**
- Giant planet **300x**
- Brown dwarf **5,000x**
- Dwarf star **30,000x**

◀ ▶ ⏸ 📷

Help End Game



Years: 0.0/500
Points: 0

2 / 12 bodies
Crowdedness bonus: 1.0x
Habitability bonus: 1.0x
Central star: 1.0x
Speed: 3x

Planet 1 (1.00 M_{earth})

	Name	High score
1.	Ben	51,256,990
2.	Rachael	31,585,152
3.	Ben	28,164,173
4.	Augusto	23,419,920
5.	Ben	22,753,808
6.	Augusto	20,448,158
7.	Ben	17,863,445
8.	Mike P	17,768,522
9.	Mike P	17,675,722
10.	Rachael	17,439,645

<http://www.stefanom.org/spc>

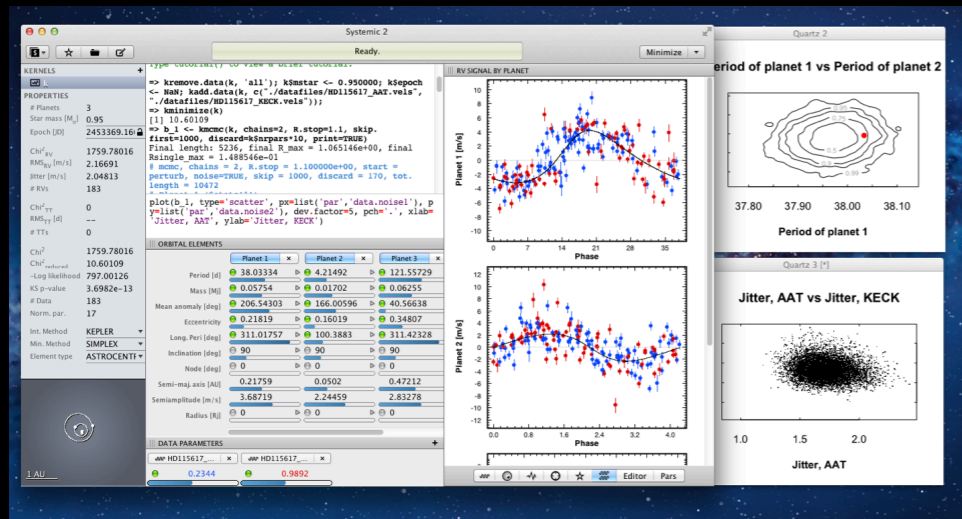
<http://www.stefanom.org/spc>

user: baesm

password: baesm

My next goal is for the game to be able to use interesting compact multi-planet systems (e.g. Kepler-11) as the starting templates, so the player can **mercilessly destabilize them by adding planets.**

Thank you!



All things Systemic:

<http://www.stefanom.org/systemic>

Play with this game and beat your fellow astronomers:

<http://www.stefanom.org/spc>

user/password: **baesm**



Here is a list of all the tools I mentioned in this talk (with links):

<https://github.com/stefano-meschiari/Notes>